

Rare Isotope Physics

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June 13th, 2024

- 1. Introduction**
- 2. Shell model & clustering**
- 3. Nuclear radioactivity**
- 4. Nuclear reactions**

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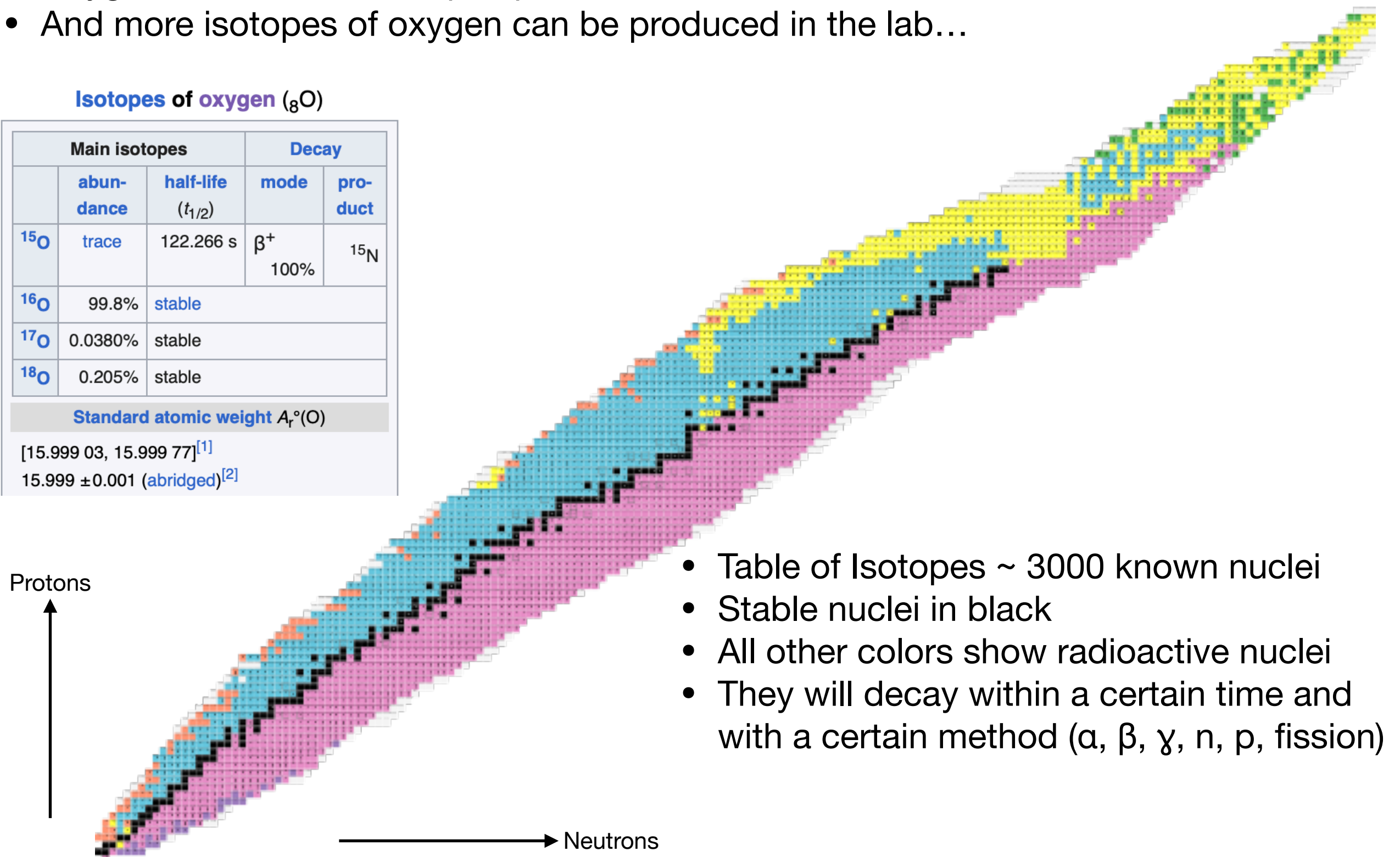
- Isotopes = Nuclei with the same number of protons
- Oxygen = 4 different isotopes present in nature: ^{15}O , ^{16}O , ^{17}O , ^{18}O
- And more isotopes of oxygen can be produced in the lab...

Isotopes of oxygen (^8O)

Main isotopes			Decay	
	abun- dance	half-life ($t_{1/2}$)	mode	pro- duct
^{15}O	trace	122.266 s	β^+ 100%	^{15}N
^{16}O	99.8%	stable		
^{17}O	0.0380%	stable		
^{18}O	0.205%	stable		

Standard atomic weight $A_r^\circ(\text{O})$

[15.999 03, 15.999 77]^[1]
 15.999 ± 0.001 (abridged)^[2]



- Table of Isotopes ~ 3000 known nuclei
- Stable nuclei in black
- All other colors show radioactive nuclei
- They will decay within a certain time and with a certain method (α , β , γ , n, p, fission)

Every nucleus is characterised by:

$${}^A_Z X_N$$

Z: Proton number (chemical element)

N: Neutron number

A=Z+N : Mass number

Isotopes: Nuclei with same Z

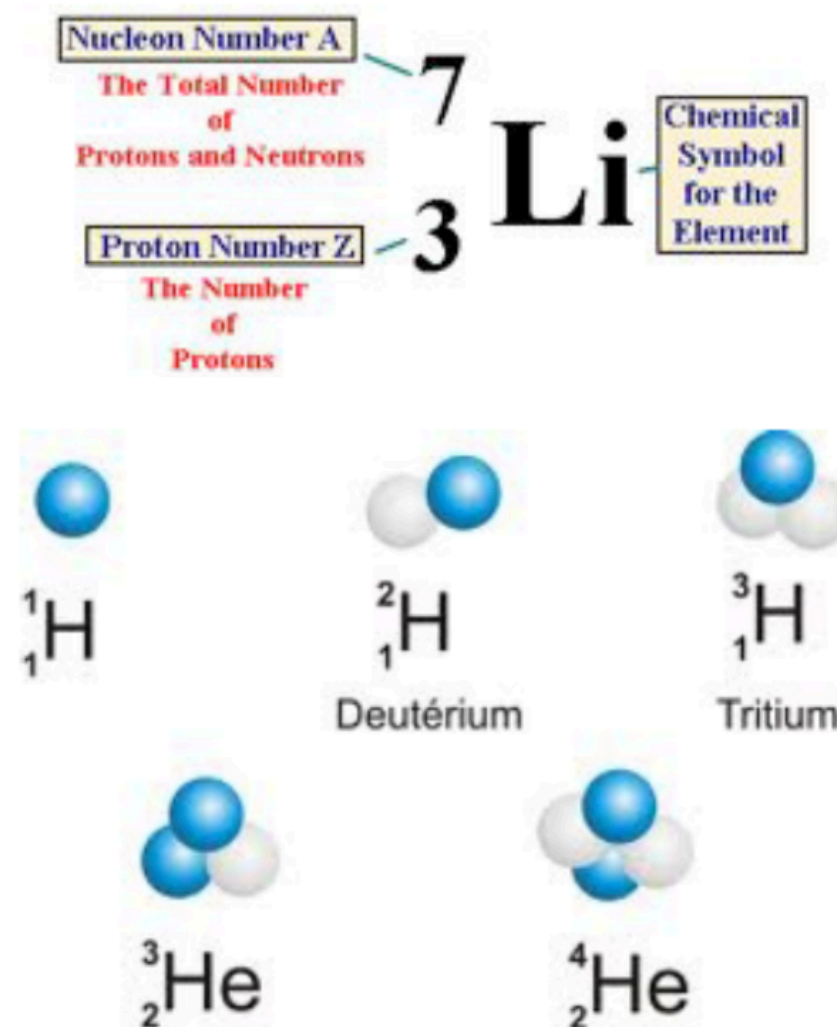
Isotones: Nuclei with same N

Isobars: Nuclei with same A

Charge of the nucleus **Ze** with $e=1.602 \cdot 10^{-19} \text{ C}$

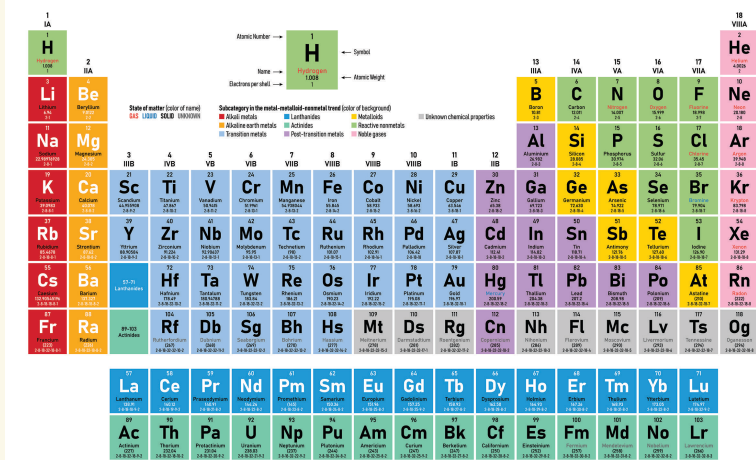
Neutral atom has Z protons and Z electrons

Ion : missing or added electron(s) ($q=1^+, 2^+, \dots, Z^+$)

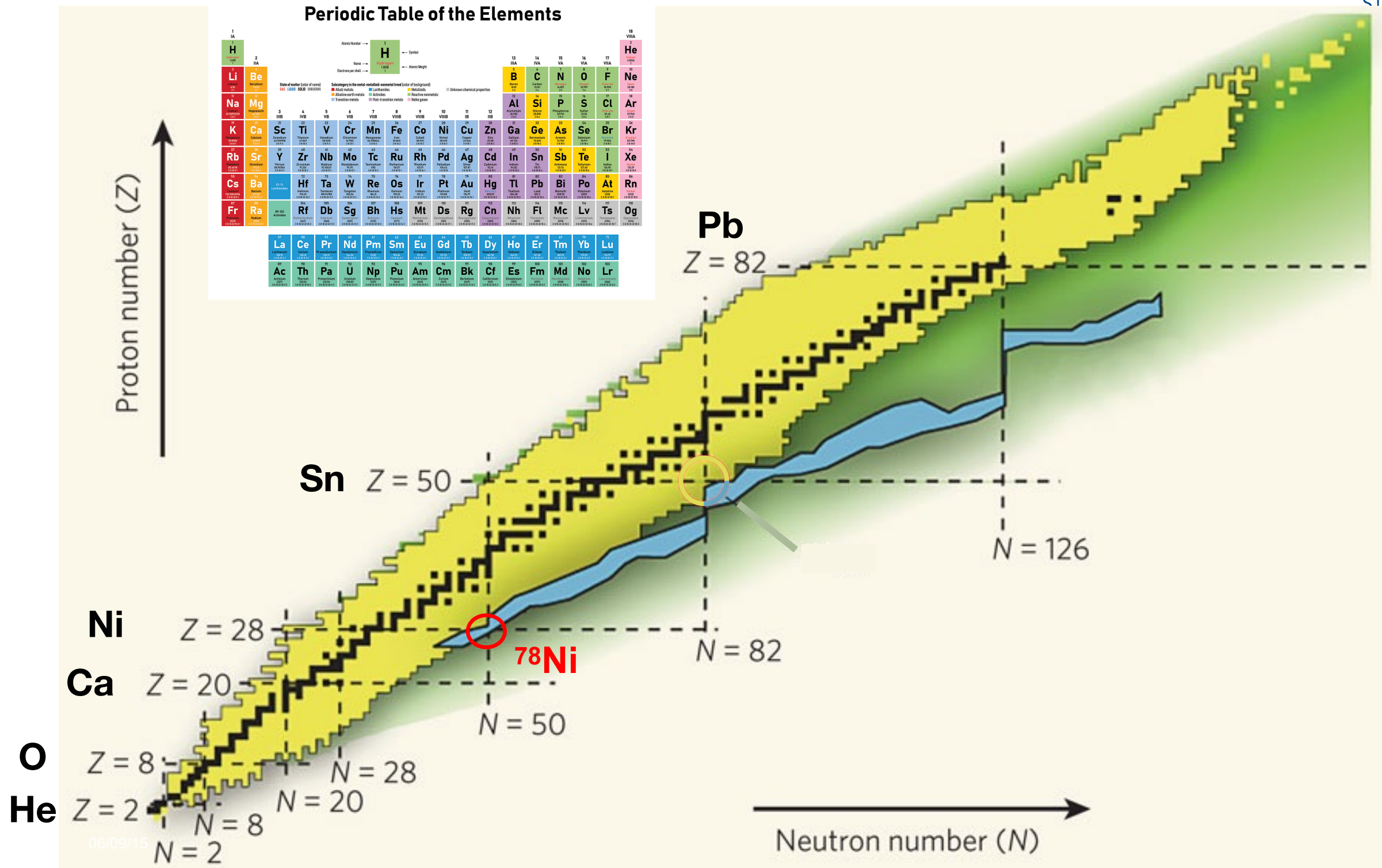


- <https://www.nndc.bnl.gov/nudat3/>

Periodic Table of the Elements

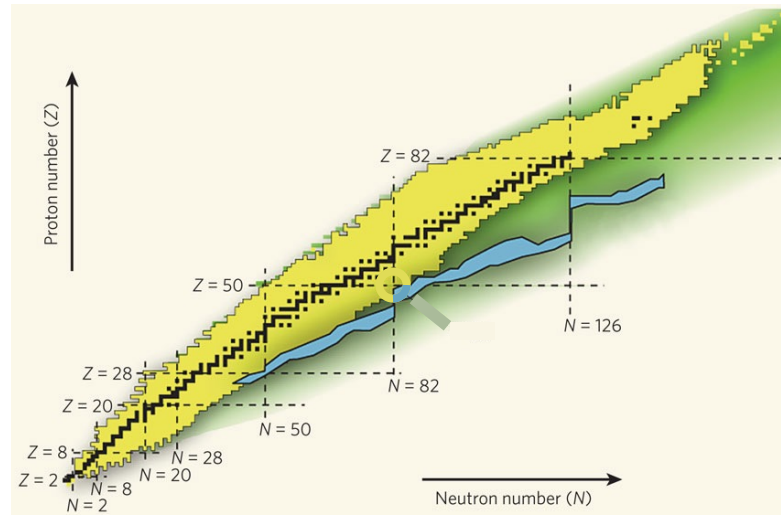


A standard periodic table of elements, color-coded by groups. It includes element symbols, atomic numbers, and names. The table is organized into rows and columns, with groups labeled at the top and bottom.



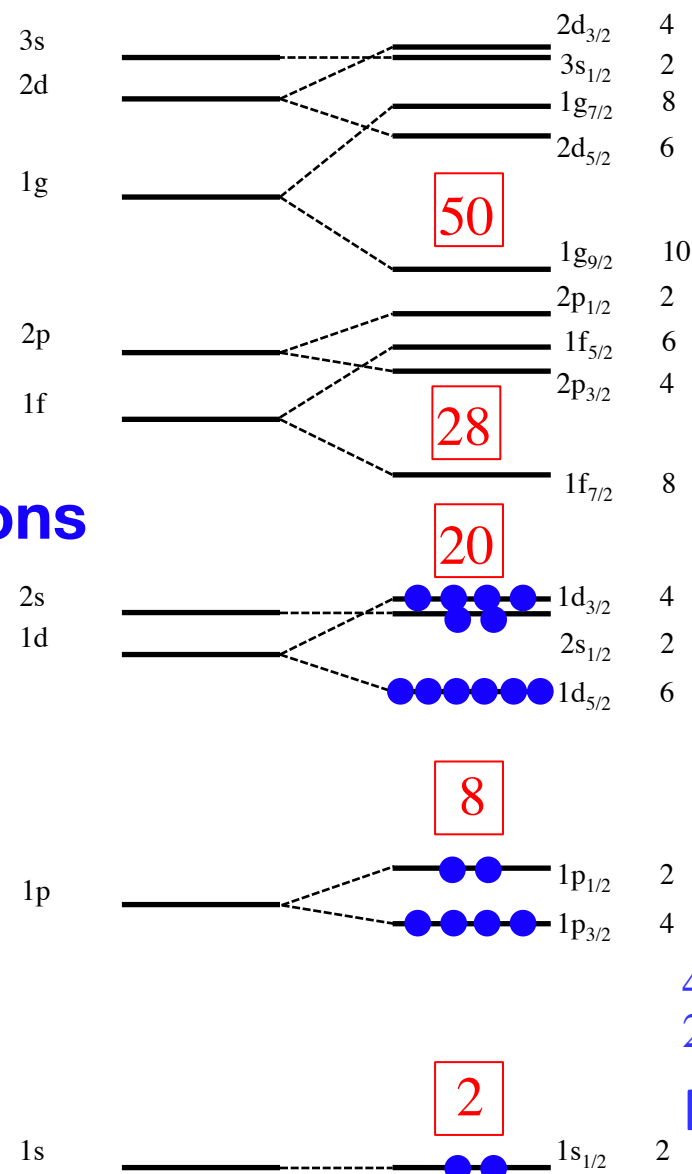
- Stable nuclei (in black) form the “valley of stability”
- Radioactive nuclei known (in yellow) and all other possible unknown (in green)

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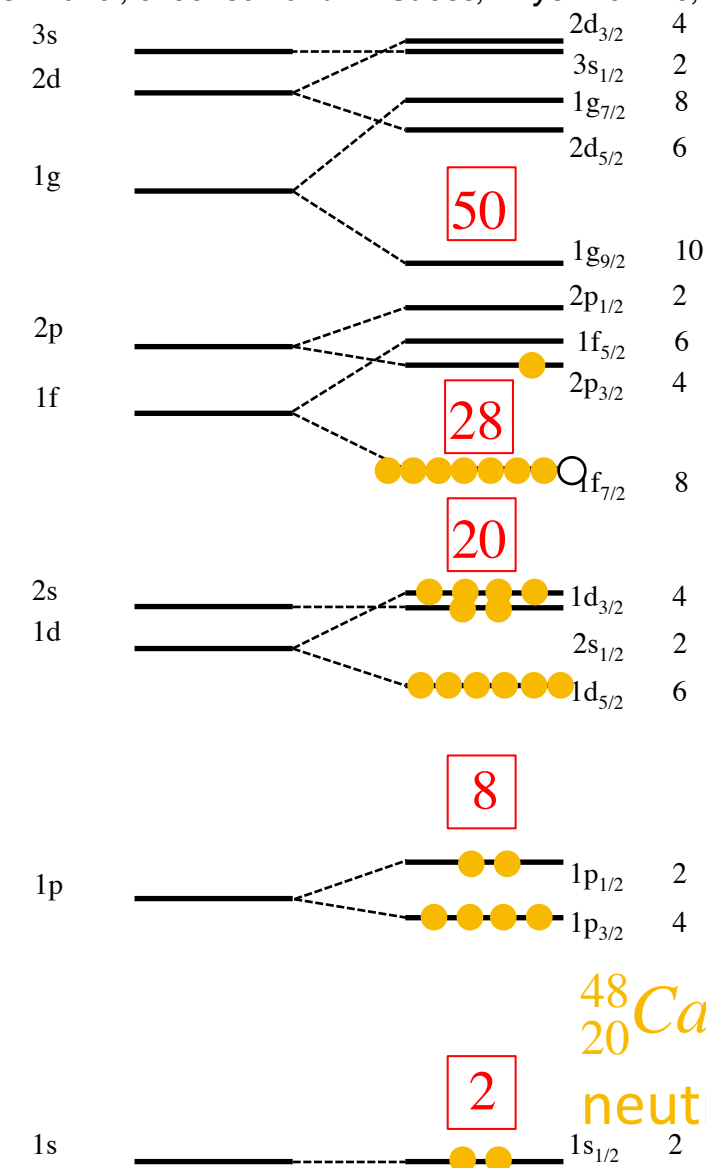
- Magic numbers= 2, 8, 20, 28, 50, 82...
- Corresponds to closed (or full) shells in our nuclear shell model
- Magic number signature= high first excited state
- Example: ^{48}Ca , doubly magic

20 protons



$^{48}_{20}\text{Ca}_{28}$
Protons

M. Goeppert-Mayer, Phys. Rev. 75, 1969 (1949).
O. Haxel, J. Jensen and H. Suess, Phys. Rev. 75, 1766 (1949).



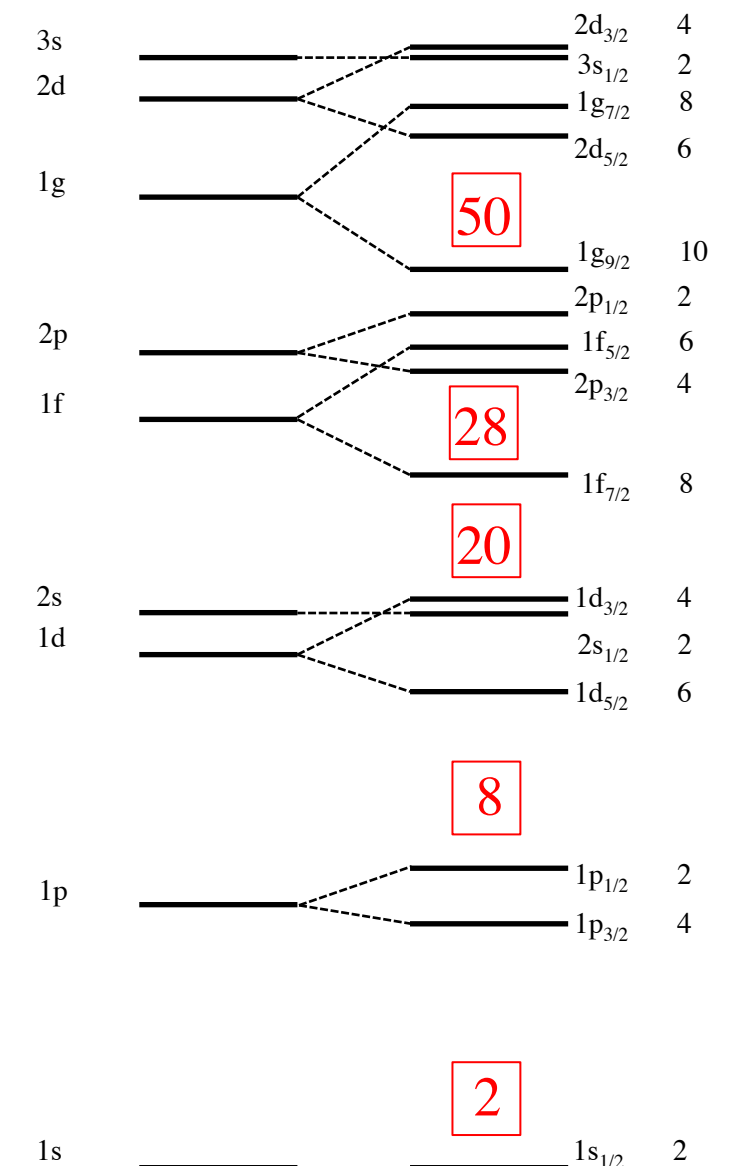
28 neutrons

$^{48}_{20}\text{Ca}_{28}$
neutrons

- Quantum numbers to define possible states:
 - Principal quantum number n
 - Orbital angular momentum ℓ ($\ell = 0, 1, 2, 3, 4$ are s, p, d, f, g orbitals)
 - Spin s (here $+1/2$ or $-1/2$, as both protons and neutrons are fermions)
 - Total angular momentum j , here $j = \ell + s$ or $\ell - s$

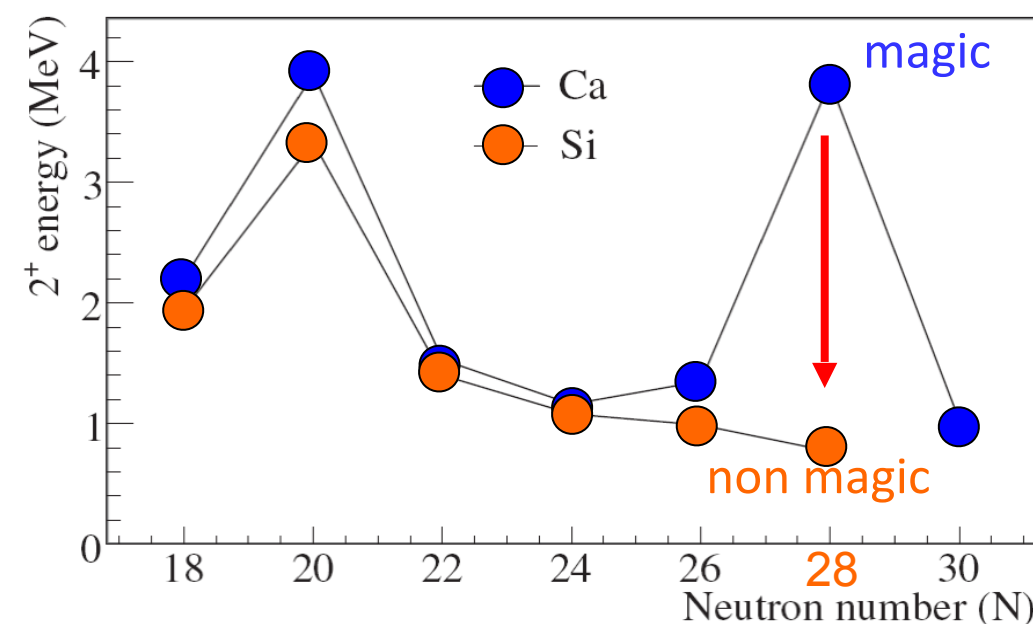
A nuclear shell is uniquely defined by : $n\ell_j$

The spin-orbit coupling is splitting the state with the same orbital angular momentum (electron shells) into TWO states now, using the total angular momentum as well

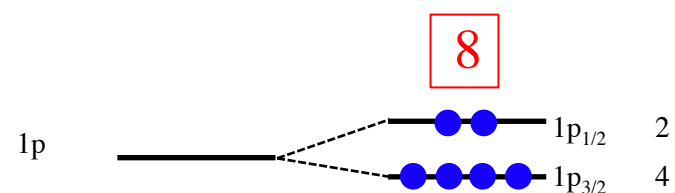
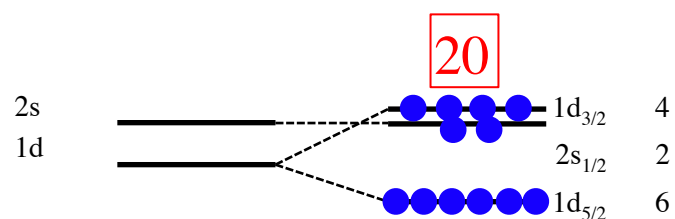


- Magic numbers= 2, 8, 20, 28, 50, 82...
- As we go to exotic nuclei, magic numbers are changing !
- Example: Ca (Z=20) vs Si (Z=14) for N=28

Magic number signature

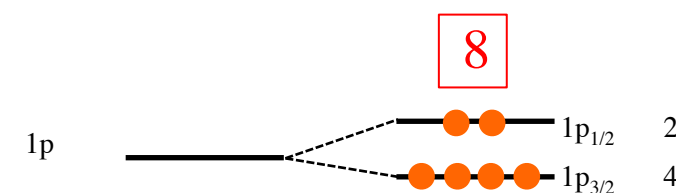
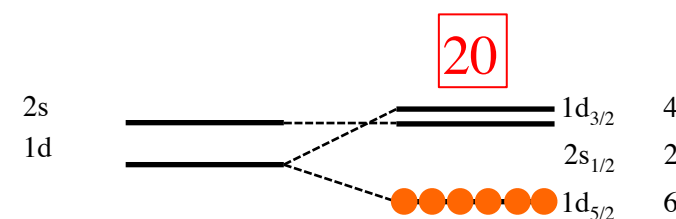


Shell structure for protons



$$E(2^+, ^{40}_{20}\text{Ca}_{20}) = 3.904\text{MeV}$$

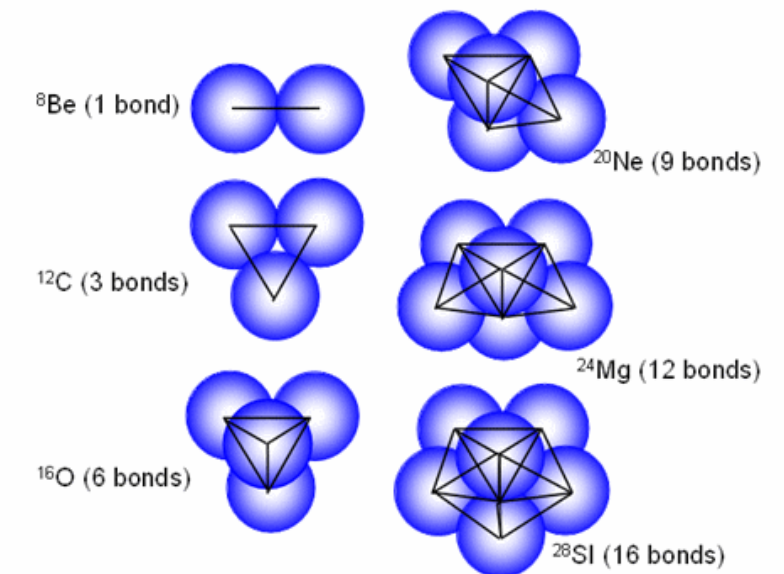
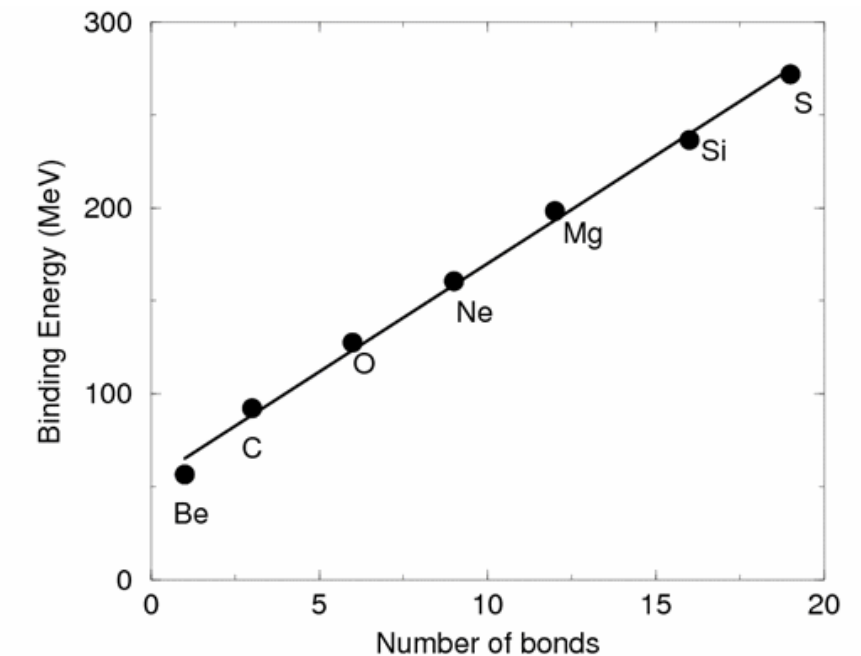
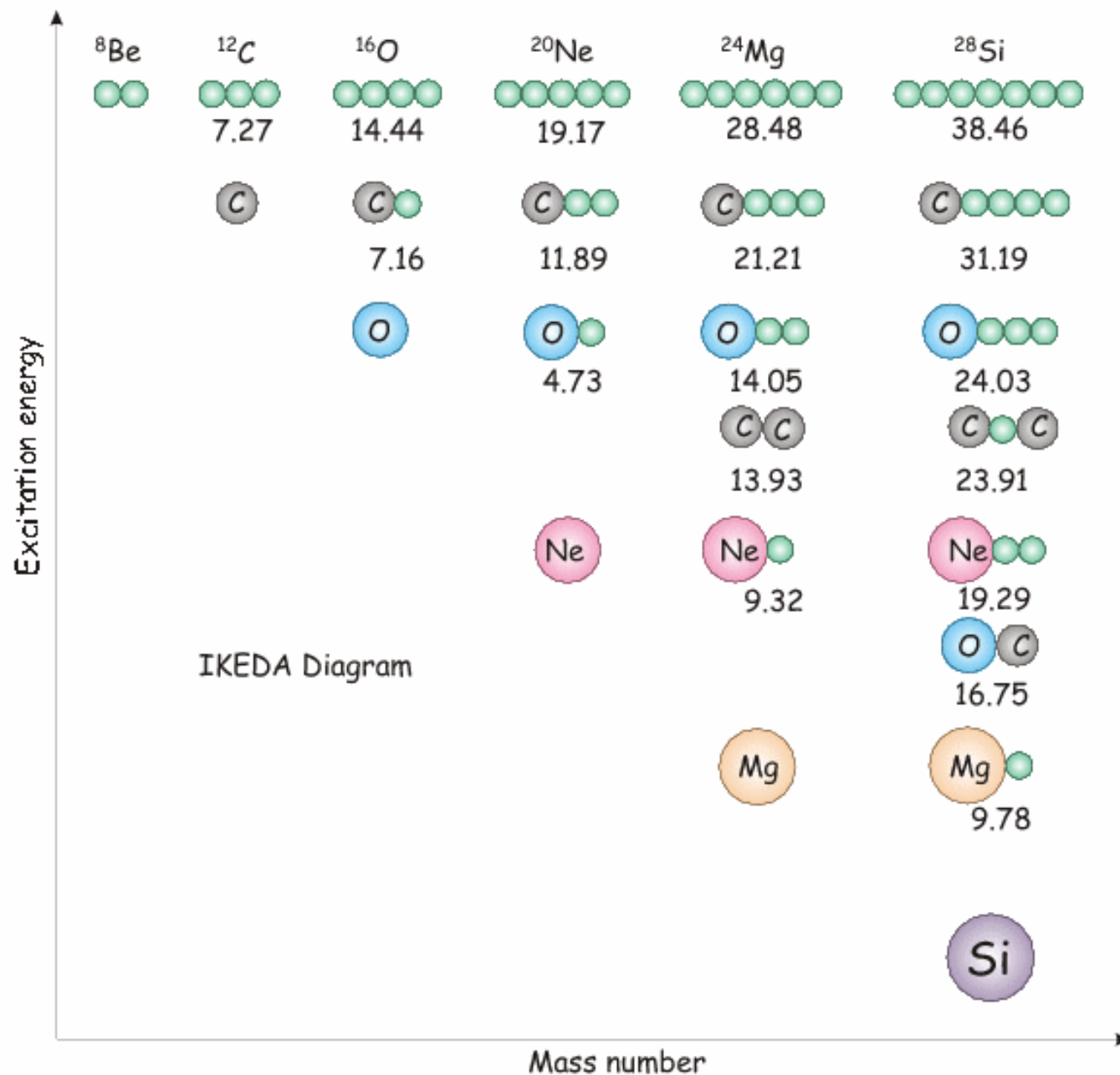
$$E(2^+, ^{48}_{20}\text{Ca}_{28}) = 3.832\text{MeV}$$



$$E(2^+, ^{34}_{14}\text{Si}_{20}) = 3.327\text{MeV}$$

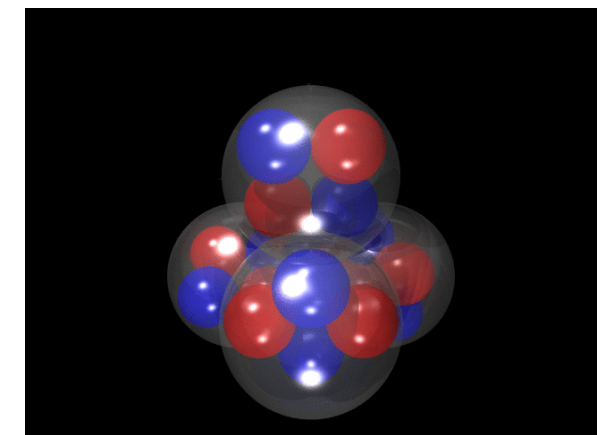
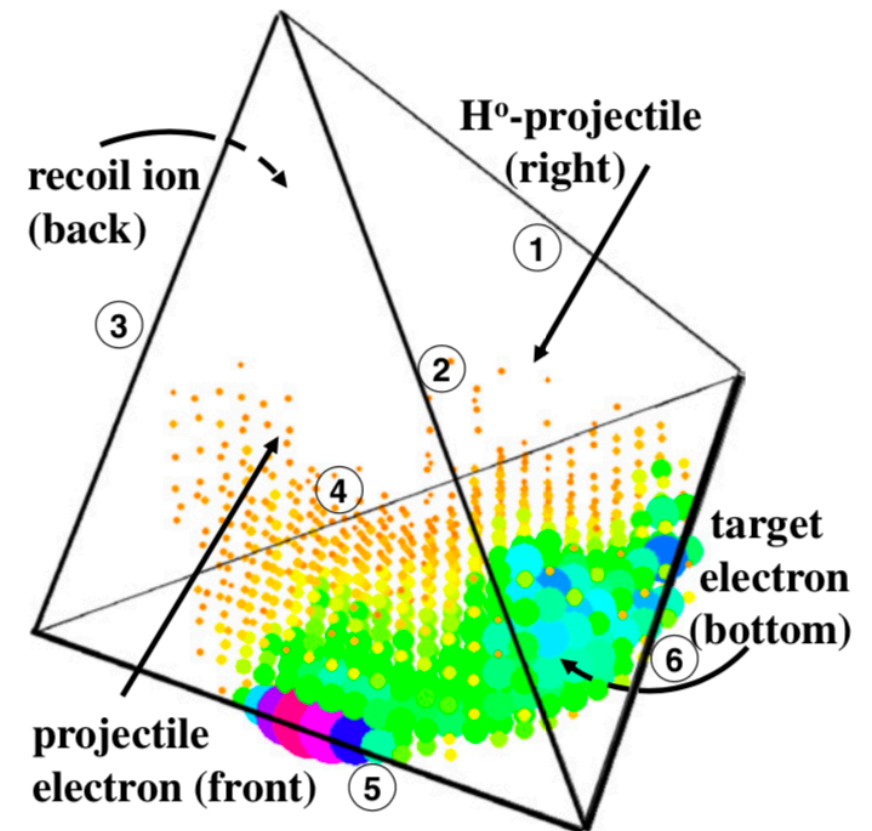
$$E(2^+, ^{42}_{14}\text{Si}_{28}) = 0.742\text{MeV}$$

With the α decay of heavy nuclei, came the idea that clusters of α particles (${}^4\text{He} = 2p + 2n$) might be preformed prior to emission

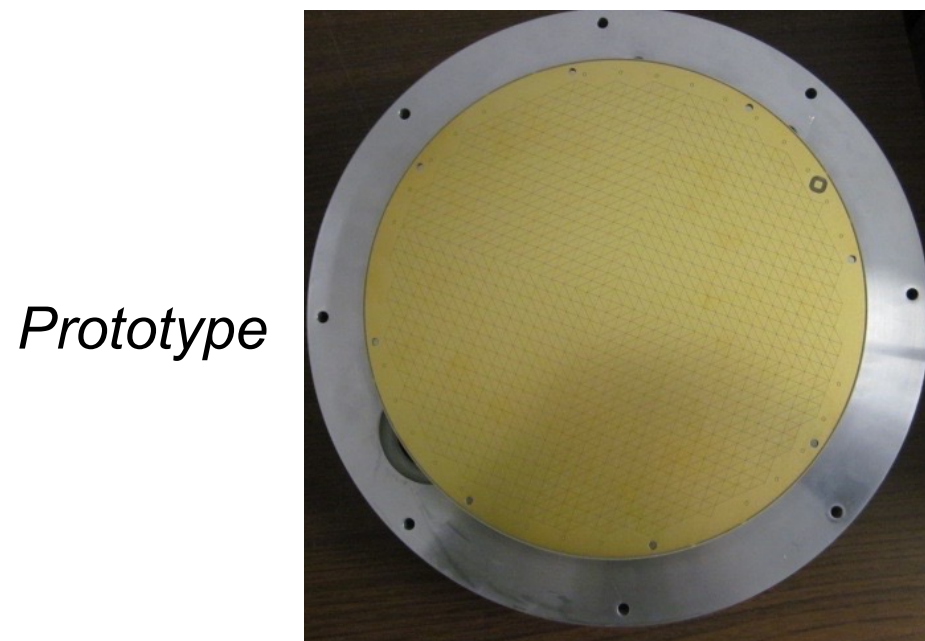
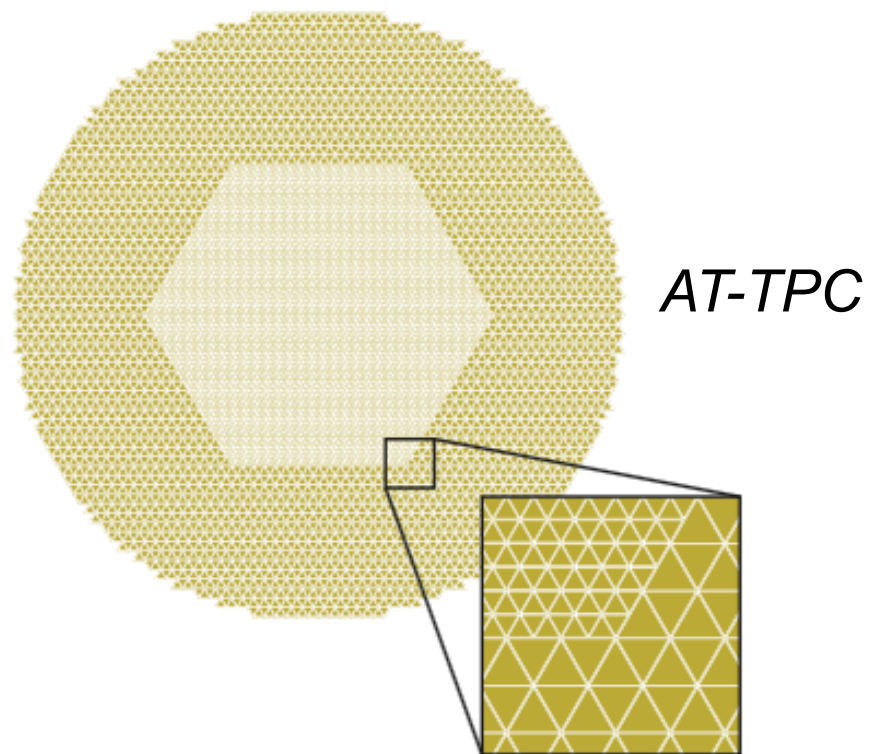


Martin Freer 2007 *Rep. Prog. Phys.* **70** 2149

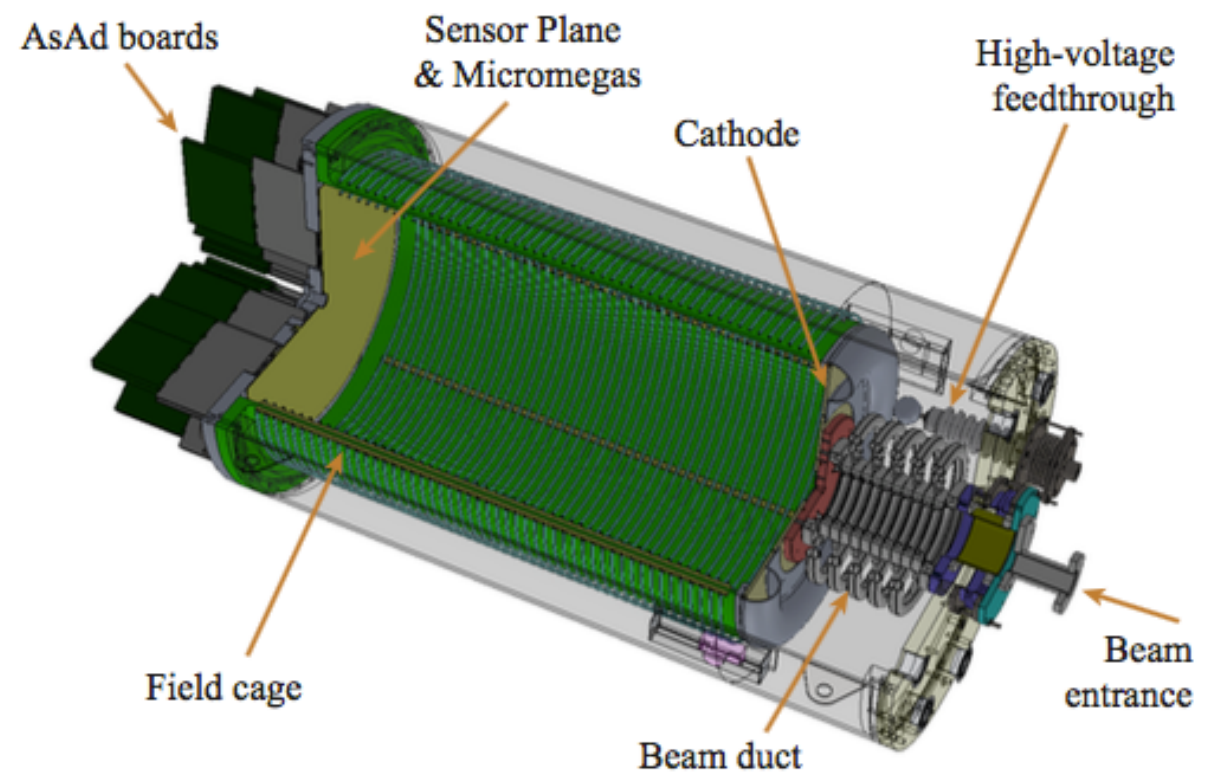
- ^4He = 2 protons + 2 neutrons, or also called α particle
- $2\alpha = ^8\text{Be}$ is UNBOUND - doesn't exist
- triple α process in stars instead to produce $^{12}\text{C} = 3\alpha$
- Hoyle state in ^{12}C = excited, spinless state that is produced in the triple α process
- ^{12}C Hoyle state confirmed to be a 3α cluster
- Mostly sequential decay = not linear chain structure
- Structure as linear chain challenged by some theories and its mostly sequential decay
- No other heavier nuclei have confirmed α -condensate states, Hoyle-like / linear chain
- Study of ^{16}O like R. Smith et al., PRL 119, 132502 (2017)



- ➡ Direct measurement of 4 α decay from the 0^+_{6} Hoyle-like state candidate at 15.097(5) MeV
- ➡ Analysis of decay with 4-particle Dalitz plot



- Active Target of 1 m length, 55 cm diameter
- ➔ Thick target, good resolution, 4π detection
- MicroMegas detection pad plane
- 10,240 pads, equilateral triangles
- GET electronics for internal trigger
- Coupling with magnetic field
- Prototype (1/2 size) used for travel with ~2000 triangular pads

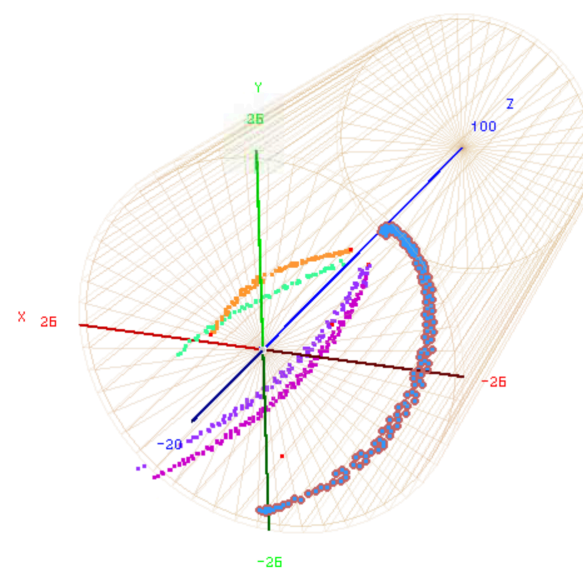
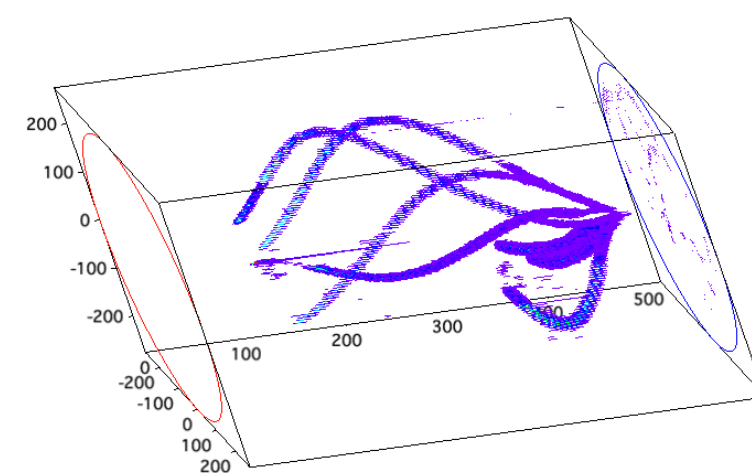
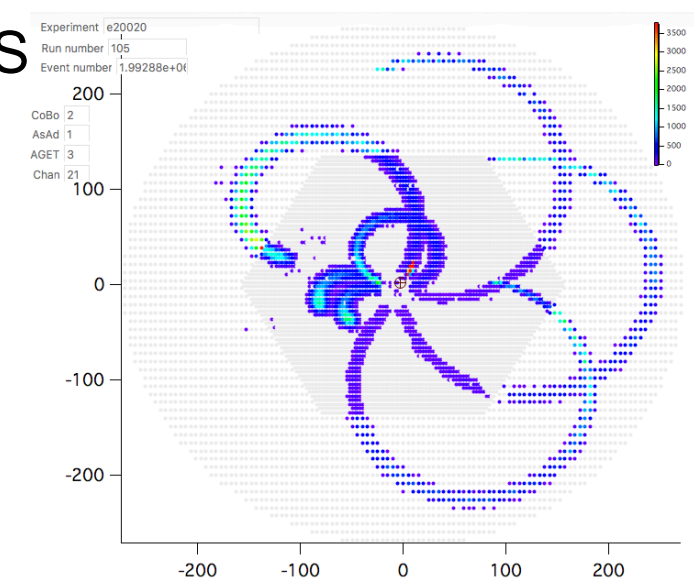


Resolution capabilities

- Scattering angle = 1° for (α, α')
- Energy resolution of 30-40 keV/u in c.m.

- ReA campaign in 2021 during transition to FRIB
- Stable or long-lived isotopes beams
- Upgrade to ReA6 with beams up to 10 AMeV
- New solenoid SOLARIS (4T) coupled to AT-TPC or HELIOS

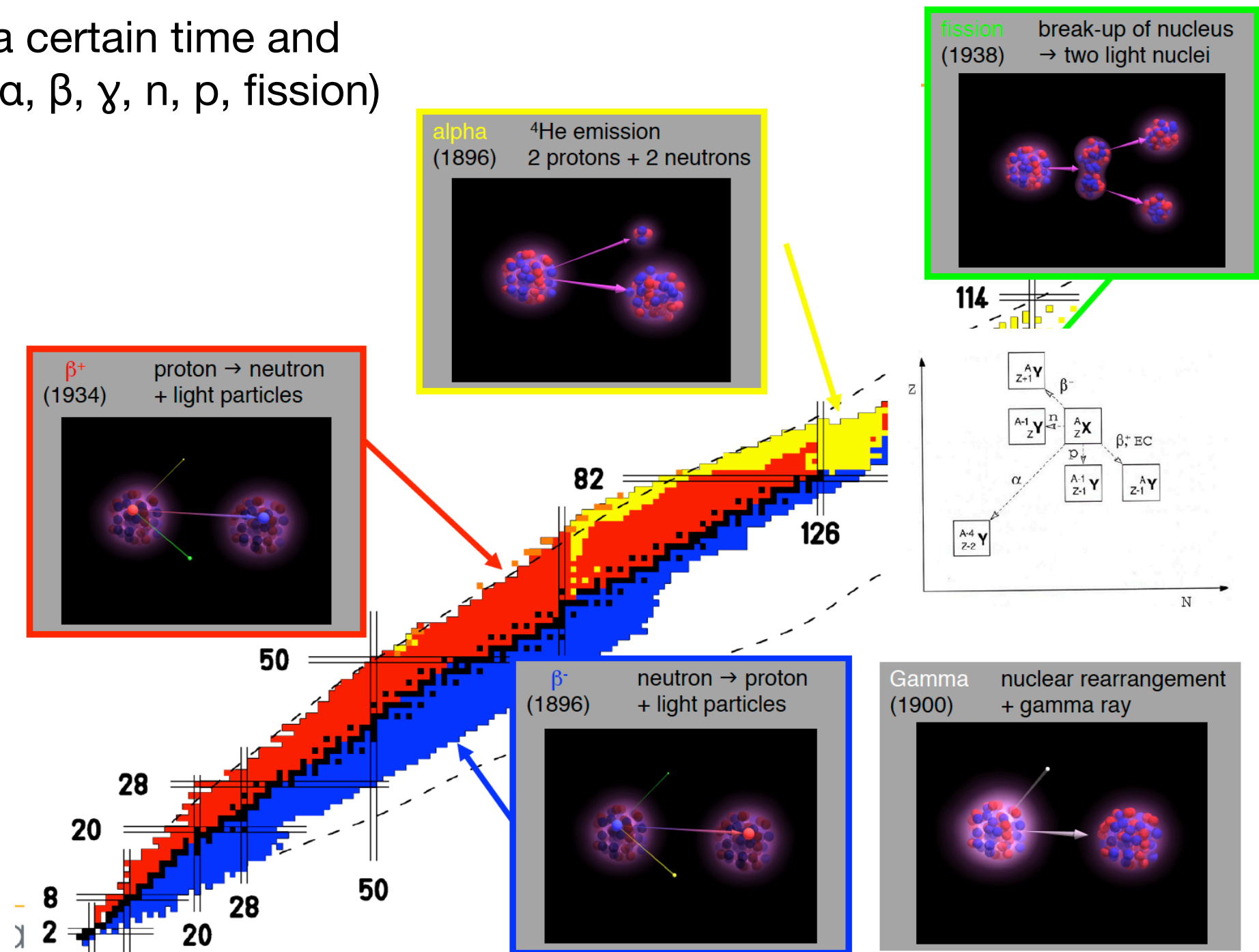
Reaction	Smith et al. ¹² C (α , α')	AT-TPC ¹⁶ O (α , α')
Energy (MeV/nucleon)	10	10
Beam Intensity (pps)	1E+10	5E+03
Beam Time (hours)	60h	120h
Target Density (at/cm ²)	5E+18	2.5E+21
Detection Efficiency	0.01	0.5
Cross section (mbarn)	0.86	0.17
Counts from N- α decay of Hoyle	93,000	465

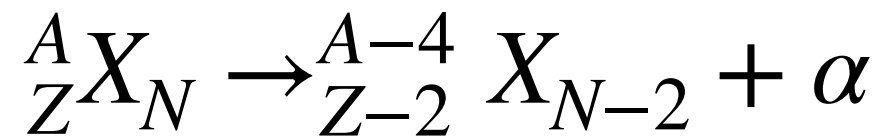


* DWUCK calculations
Direct Decay Limit (¹²C) < 0.04%

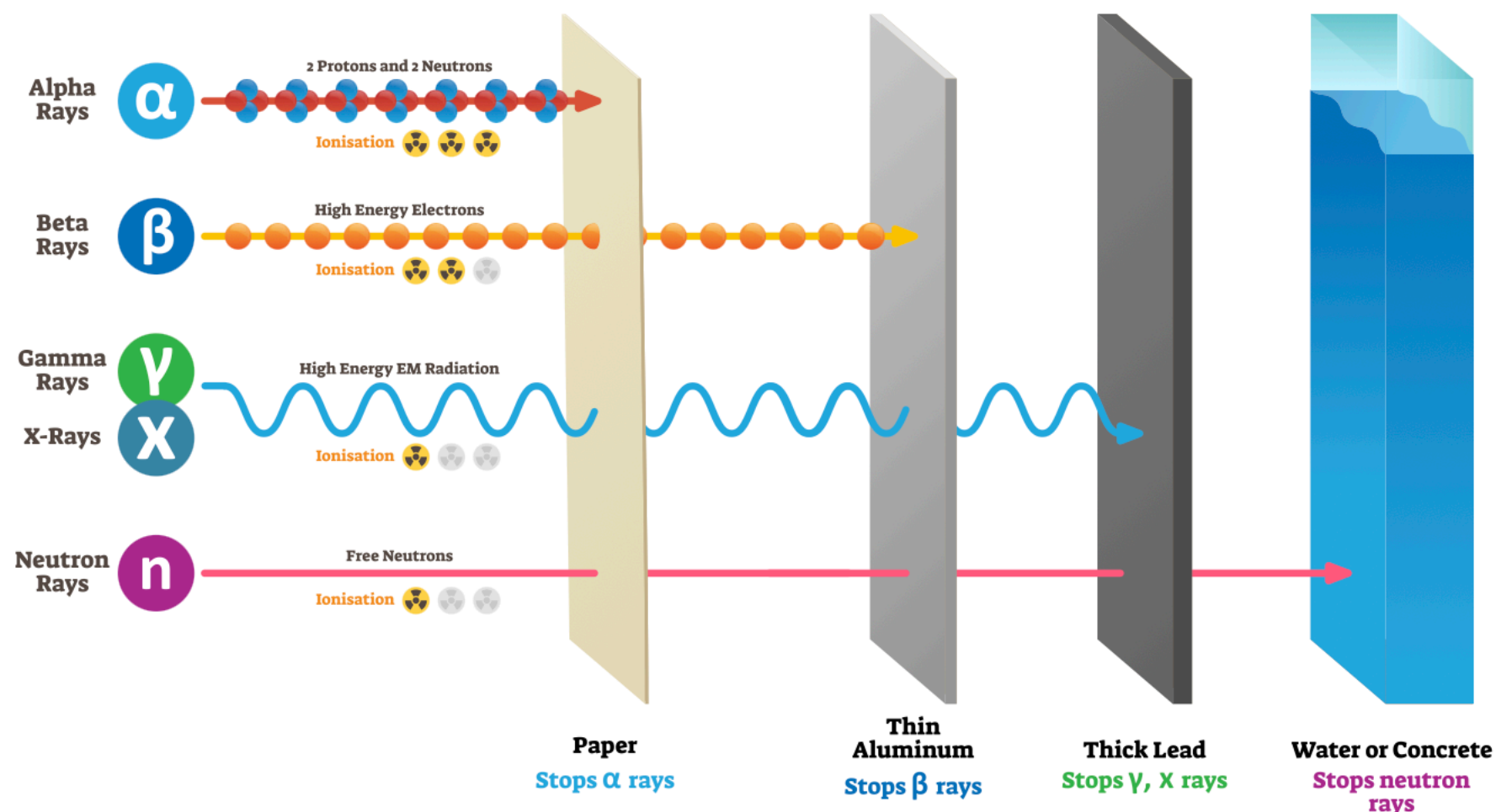
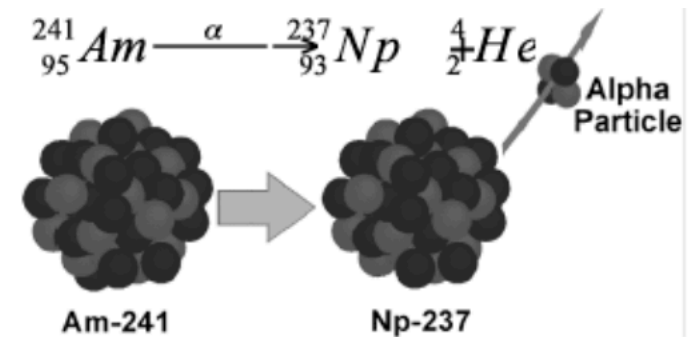
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- They will decay within a certain time and with a certain method (α , β , γ , n , p , fission)





- Particle α = nucleus ${}^4_2\text{He}$
- Possible from $Z > 50$ and $N < N_{\text{stable}}$
- Dominant for $Z > 82$



Stopped by paper and dead layer of skin

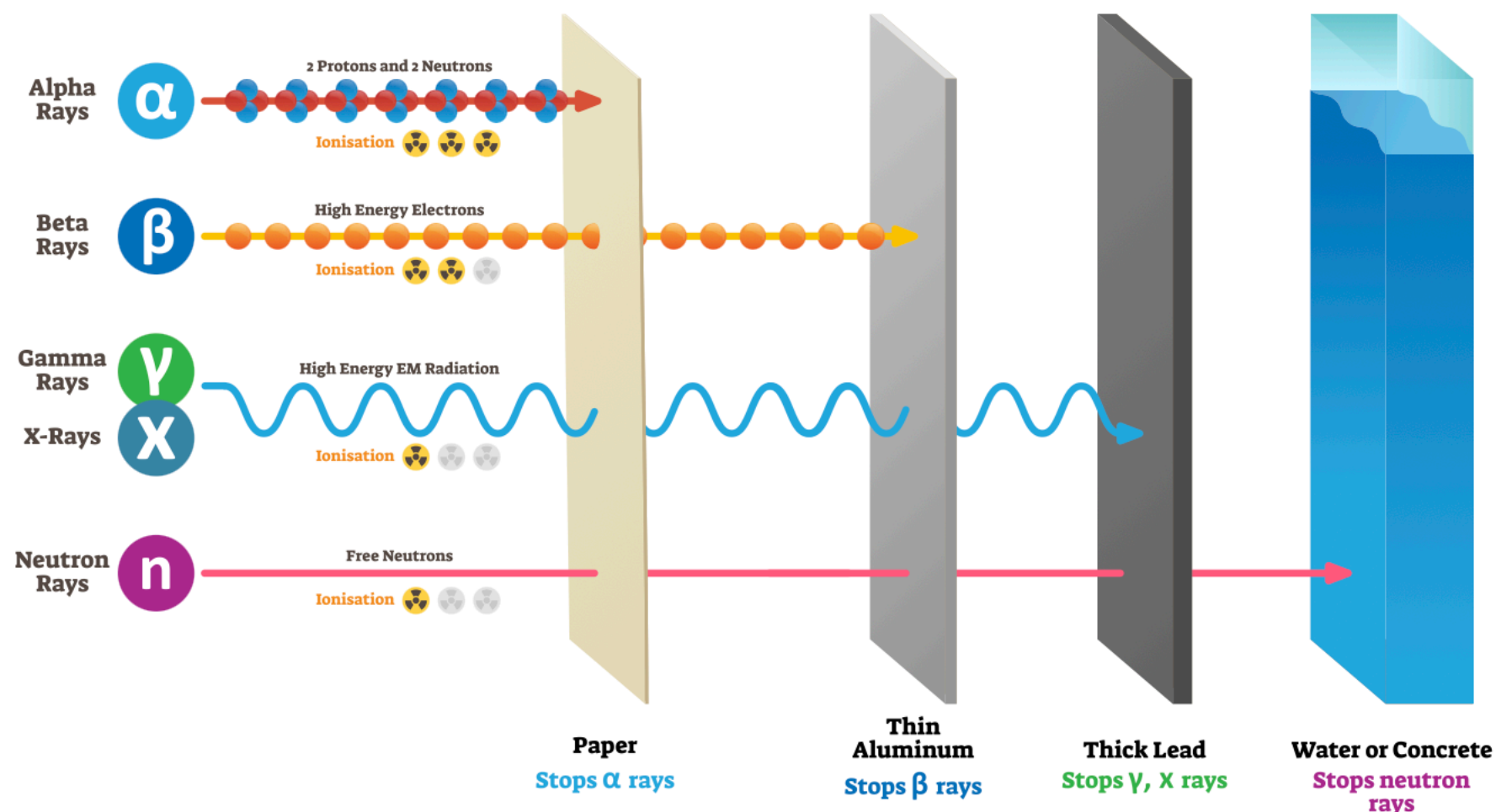
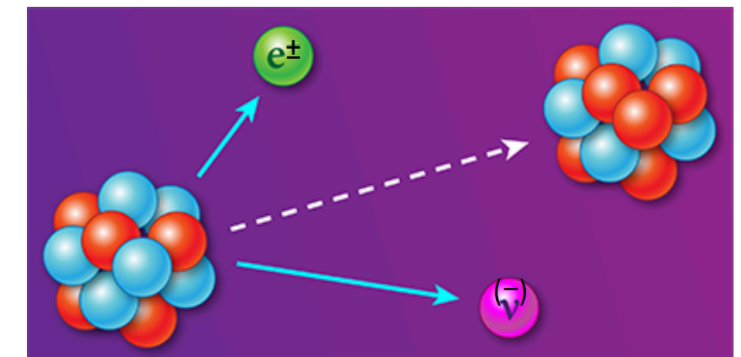
Neutron decay: $n \rightarrow p + e^- + \bar{\nu}_e$

β^- decay: ${}_Z^A X_N \rightarrow {}_{Z+1}^A Y_{N-1} + e^- + \bar{\nu}_e$ **Neutron-rich side**

β^+ decay: ${}_Z^A X_N \rightarrow {}_{Z-1}^A Y_{N+1} + e^+ + \nu_e$

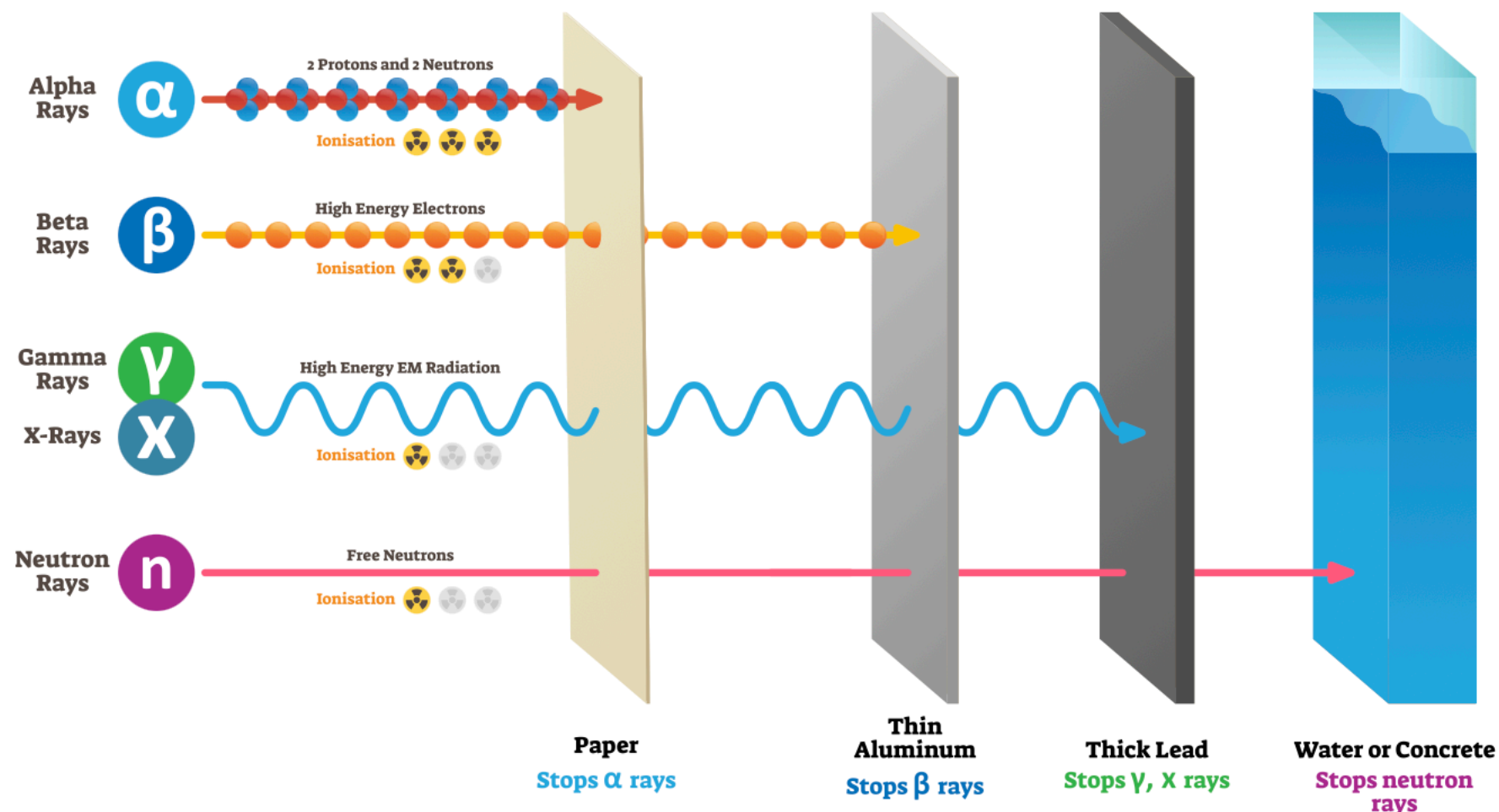
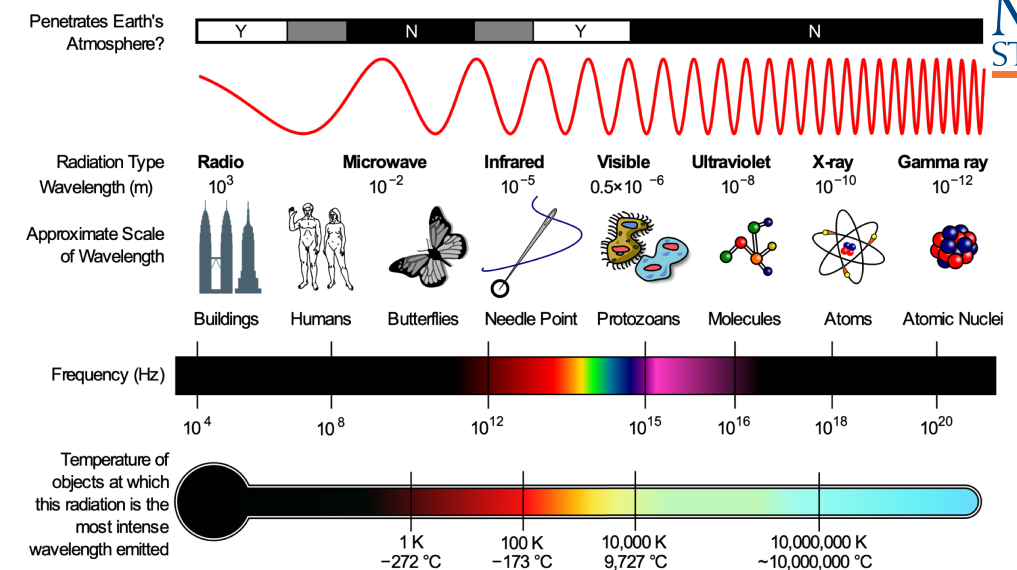
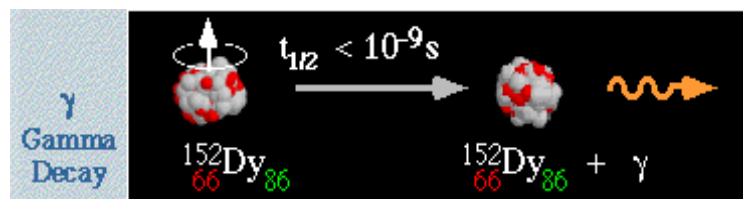
EC decay: ${}_Z^A X_N + e^- \rightarrow {}_{Z-1}^A Y_{N+1} + \nu_e$

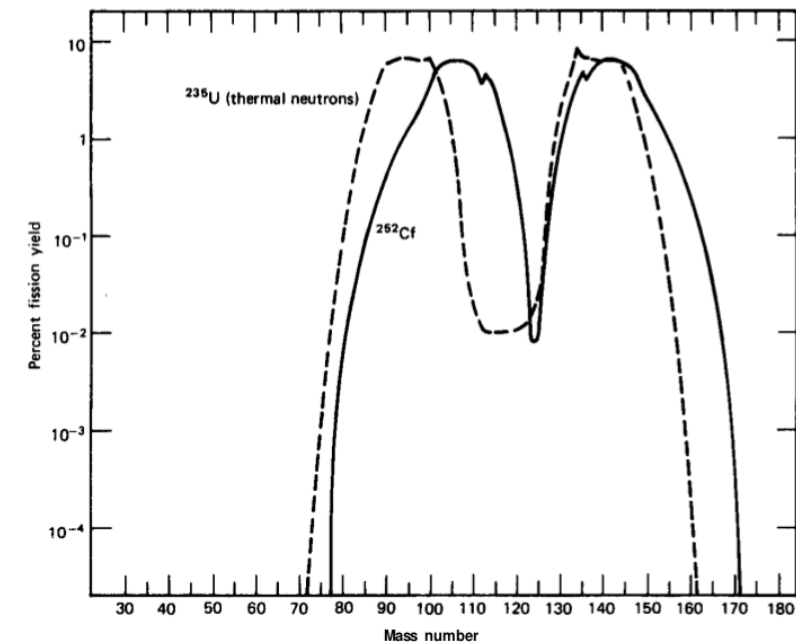
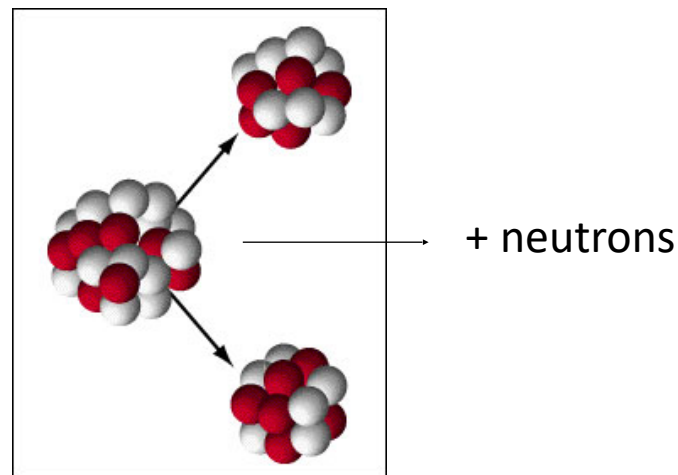
Proton-rich side



A 1 MeV beta can travel up to 4 m in air and 1 cm in plastic

- X-rays and gamma (γ) rays are photons – no charge
- A nucleus in an excited state (one nucleon in a higher shell) will want to deexcite and produce γ rays





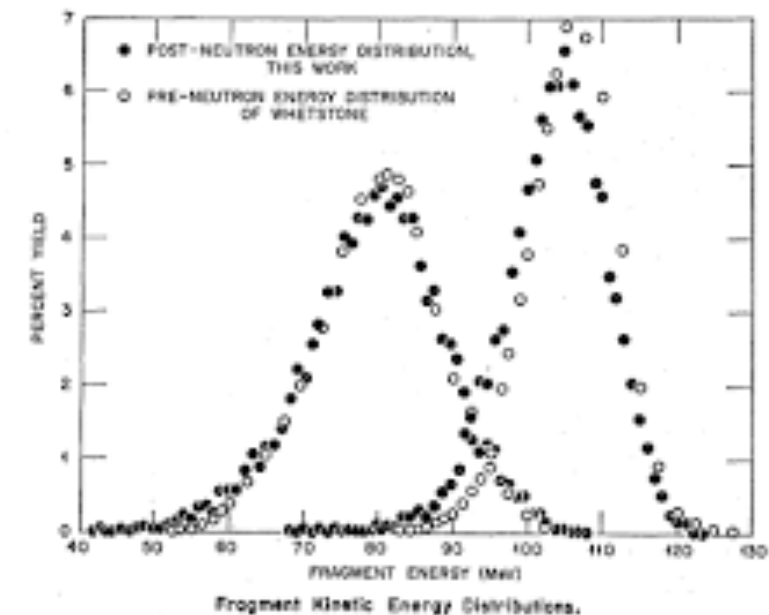
Symmetric fission

Nuclei that can spontaneously decay into two lighter fragments

Example: ^{252}Cf (half-life = 85 y if only via fission)

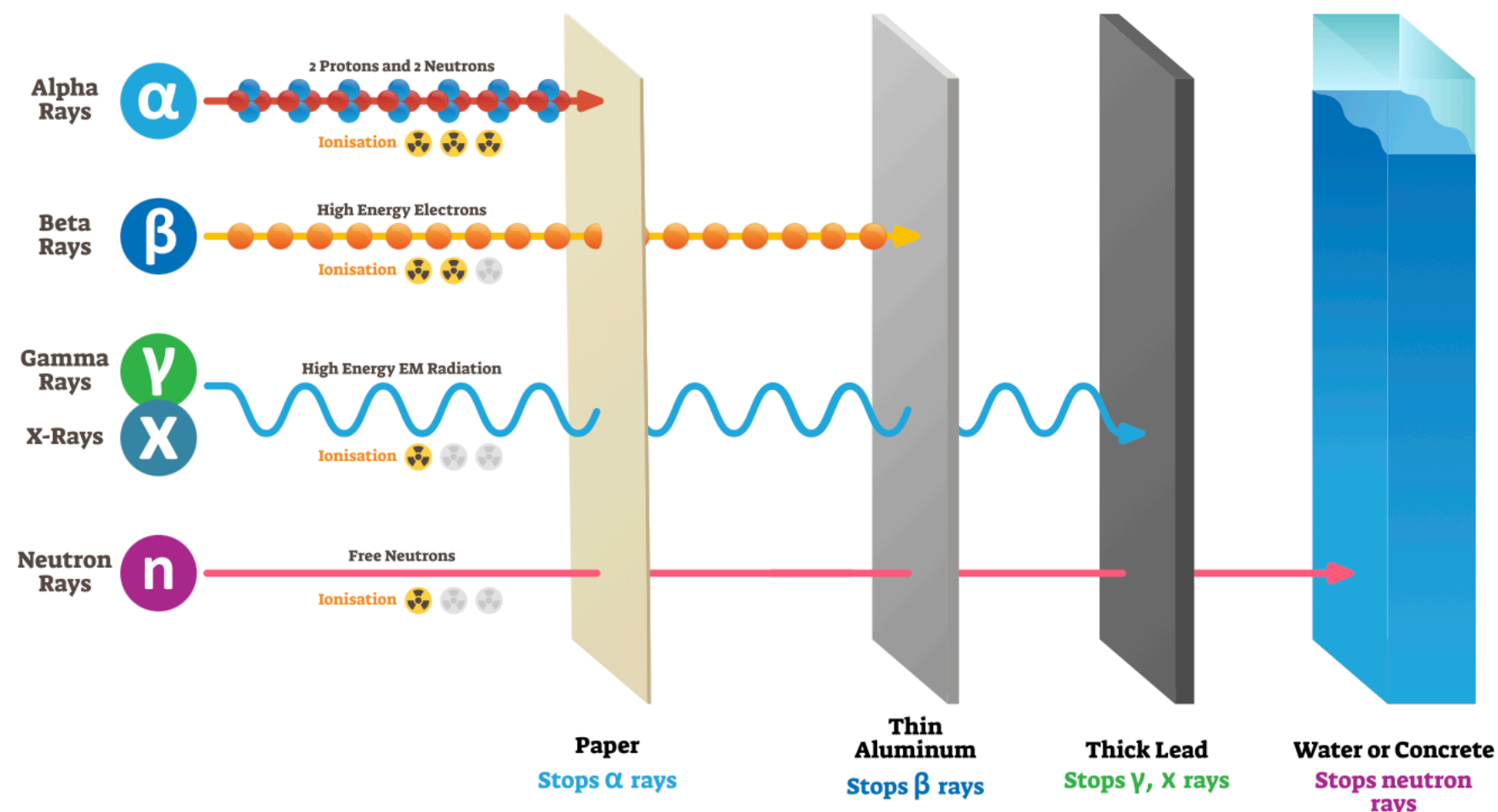
Possible when the nucleus gets big ($Z > 100$)

However, we usually have other decay processes that will compete with fission, such as alpha decay
Half-life of $^{252}\text{Cf} = 2.65\text{y}$



Asymmetric fission

- Proton decay = for very proton-rich nuclei, that have so many protons that it is not stable anymore and just rejects it
- Neutron decay = same on the neutron-rich side



Neutron shielding thickness depends on the energy of the neutron

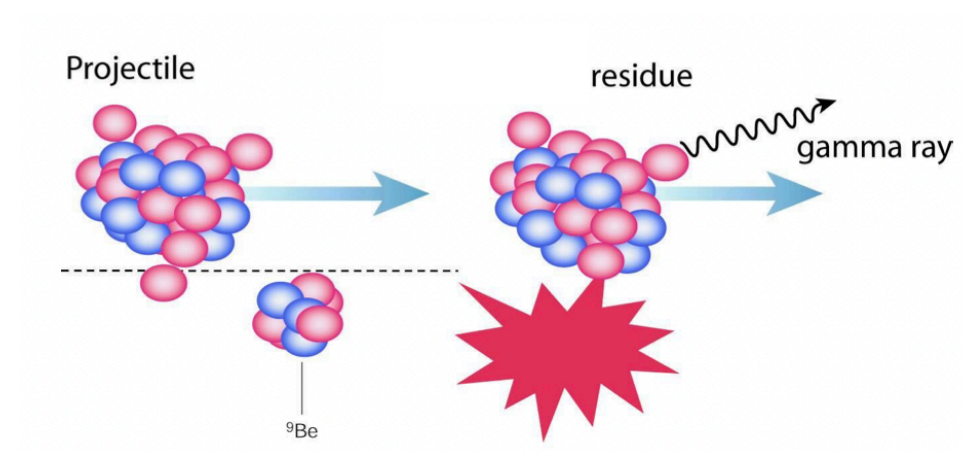
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- Most nuclei are unstable, and decay within a certain time, or half-life
- For nuclei farther away from the stable ones, their half-life gets shorter

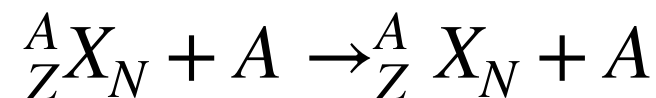
How to study nuclei that live for less than a second?

We can use the normal decay of an isotope, but then we have no control on when it happens

- ➡ Create the nucleus of interest via an induced reaction
- ➡ Use collisions and study the reaction fragments with detectors

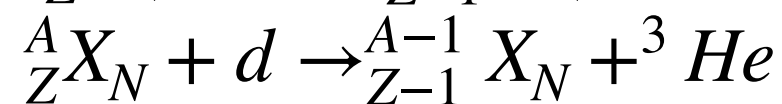
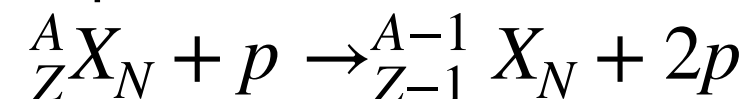


- Elastic or Inelastic collision: SAME nuclei incoming as outgoing, change in trajectory of the particle, transfer of energy (only for inelastic) and heavier nucleus at excited state

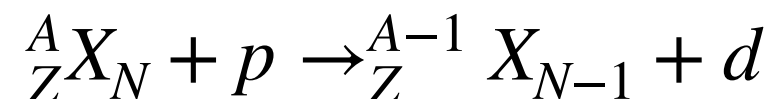


- Transfer (low energy <50 MeV/u) or knockout reactions (higher energy): remove or add one or two nucleons to a nucleus

One proton removal reactions:



One neutron removal



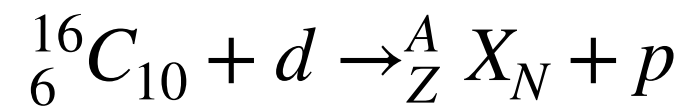
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- Fragmentation: removal of a larger number of nucleons
- Fission: breakup of a heavy nucleus into two smaller fragments
- Fusion: fusion of two nuclei into heavier one

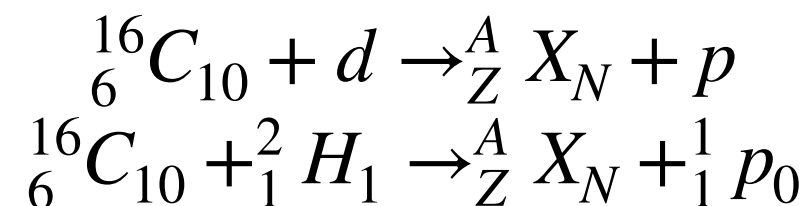
In nuclear reactions, a few principles are ALWAYS observed:

- Conservation of energy in the system = total energy in the initial system is the same in the final system
- Conservation of proton, neutron, mass number = valid for all nuclear reactions EXCEPT β decay (weak interaction)

Example: a (d, p) reaction



- (1) How many protons /neutrons are transferred? $d \rightarrow p$ so one neutron is picked up by a nucleus
- (2) What is the outgoing nucleus ${}^A_Z\text{X}_N$? Write up the conservation of proton and neutron numbers:

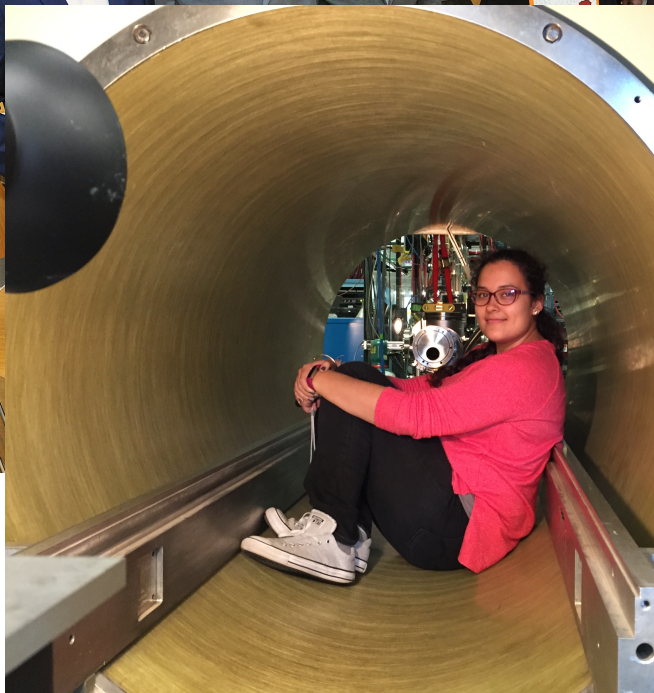
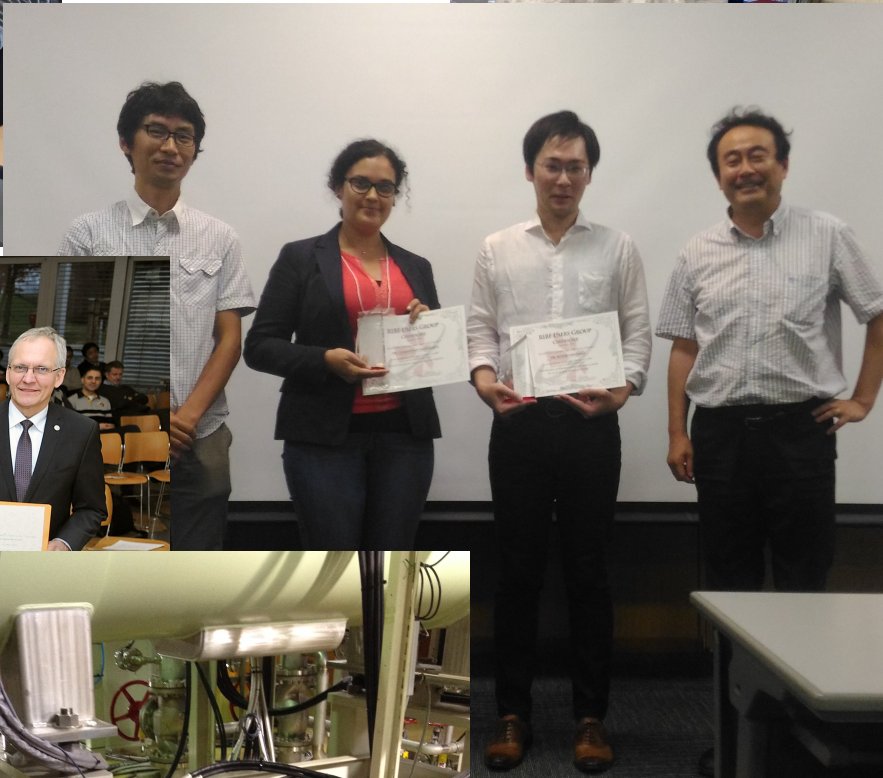


Initial	Final
$Z_{\text{initial}} = 6 + 1 = 7 = Z_{\text{tot}}$	$Z_{\text{tot}} = Z_{\text{final}} = 7 = Z + 1$
$N_{\text{initial}} = 10 + 1 = 11 = N_{\text{tot}}$	$N_{\text{tot}} = N_{\text{final}} = 11 = N + 0$

$${}^A_Z\text{X}_N$$

$$Z = 6$$

$$N = 11$$



FRIB / Michigan State University

D. Bazin*, M. Cortesi, A. Davis, W. Mittig*

RCNP & Osaka University

N. Aoi, T. Furuno*, E. Ideguchi, T. Kawabata*, A. Tamii, D. T. Tran

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J. Chen* (ANL)

J. Lou* (Peking University)

A. O. Macchiavelli* (Oakridge National Laboratory)

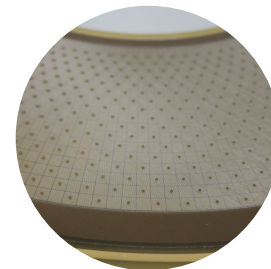
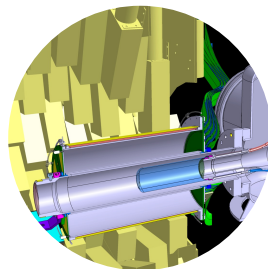
Hooi Jin Ong* (Institute of Modern Physics)

C. Santamaria (MSU²)

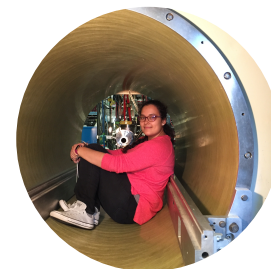
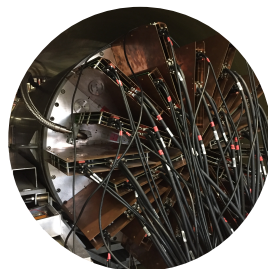
D. Suzuki* (RIKEN Nishina Center)

And the rest of the AT-TPC collaboration



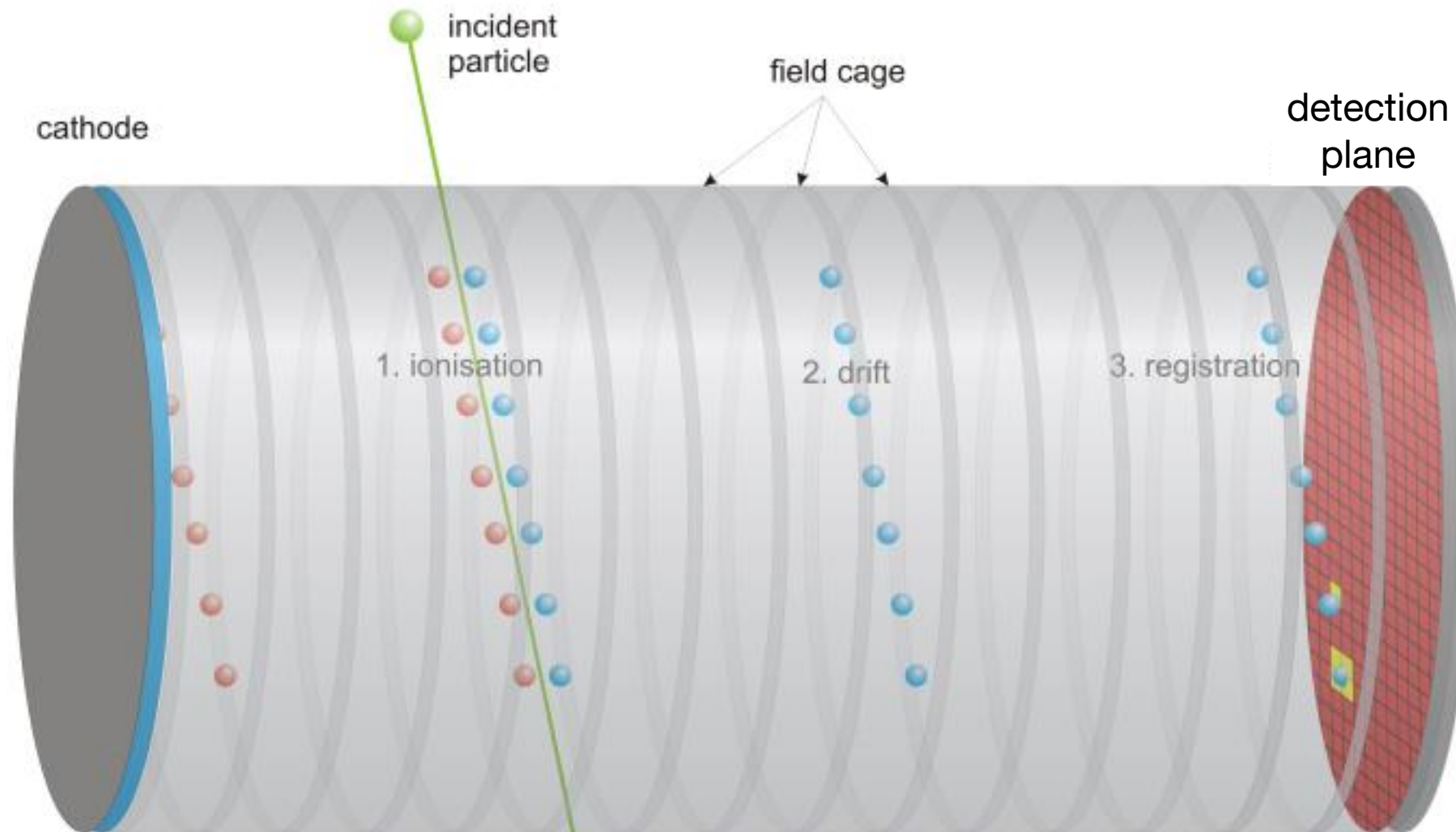


THANK YOU FOR YOUR ATTENTION !



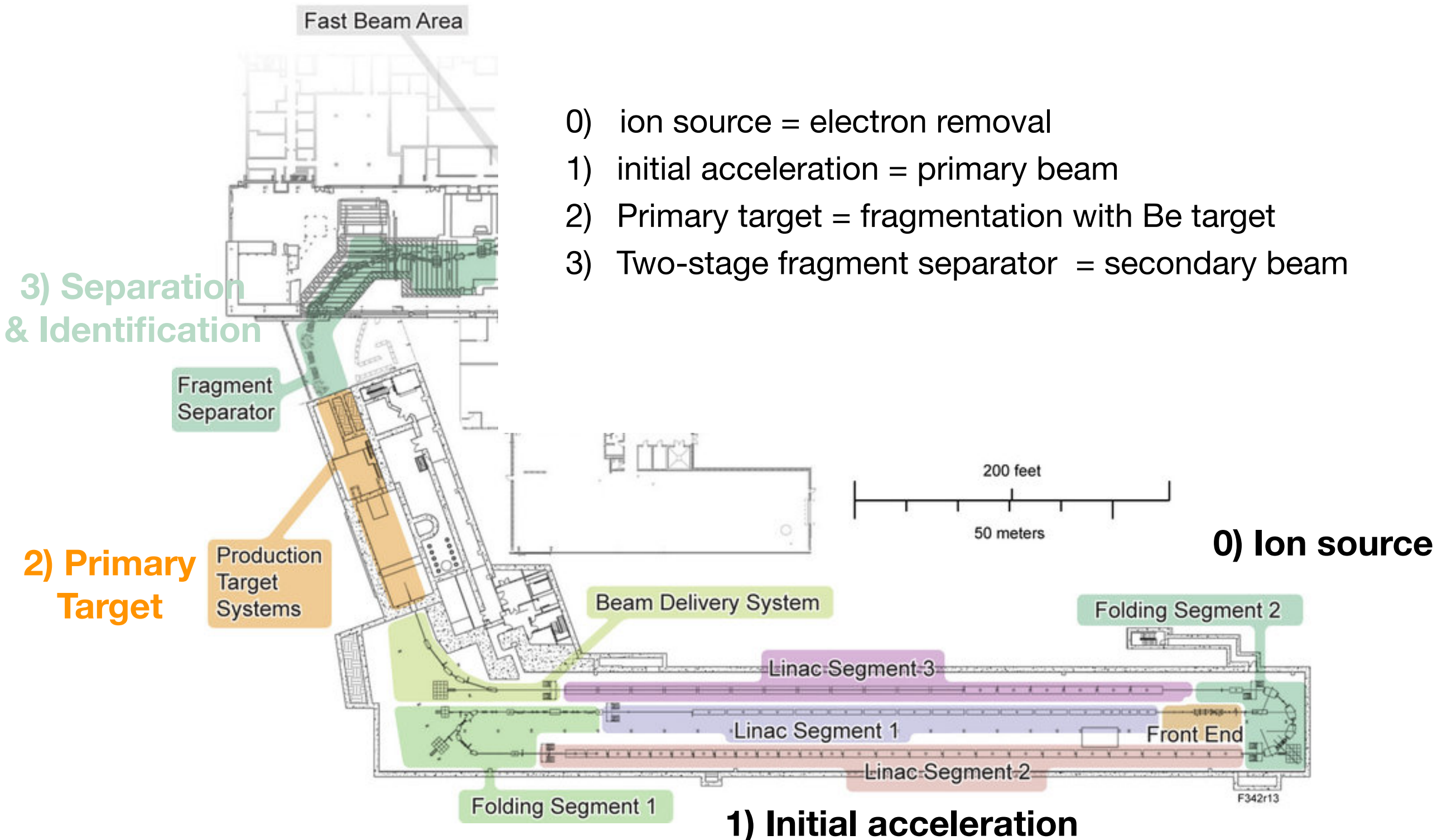
EXTRAS

- Active Target = our target material is ALSO our detector
- Mostly gaseous or liquid detectors
- Example: Time Projection Chamber (TPC) for charged particles

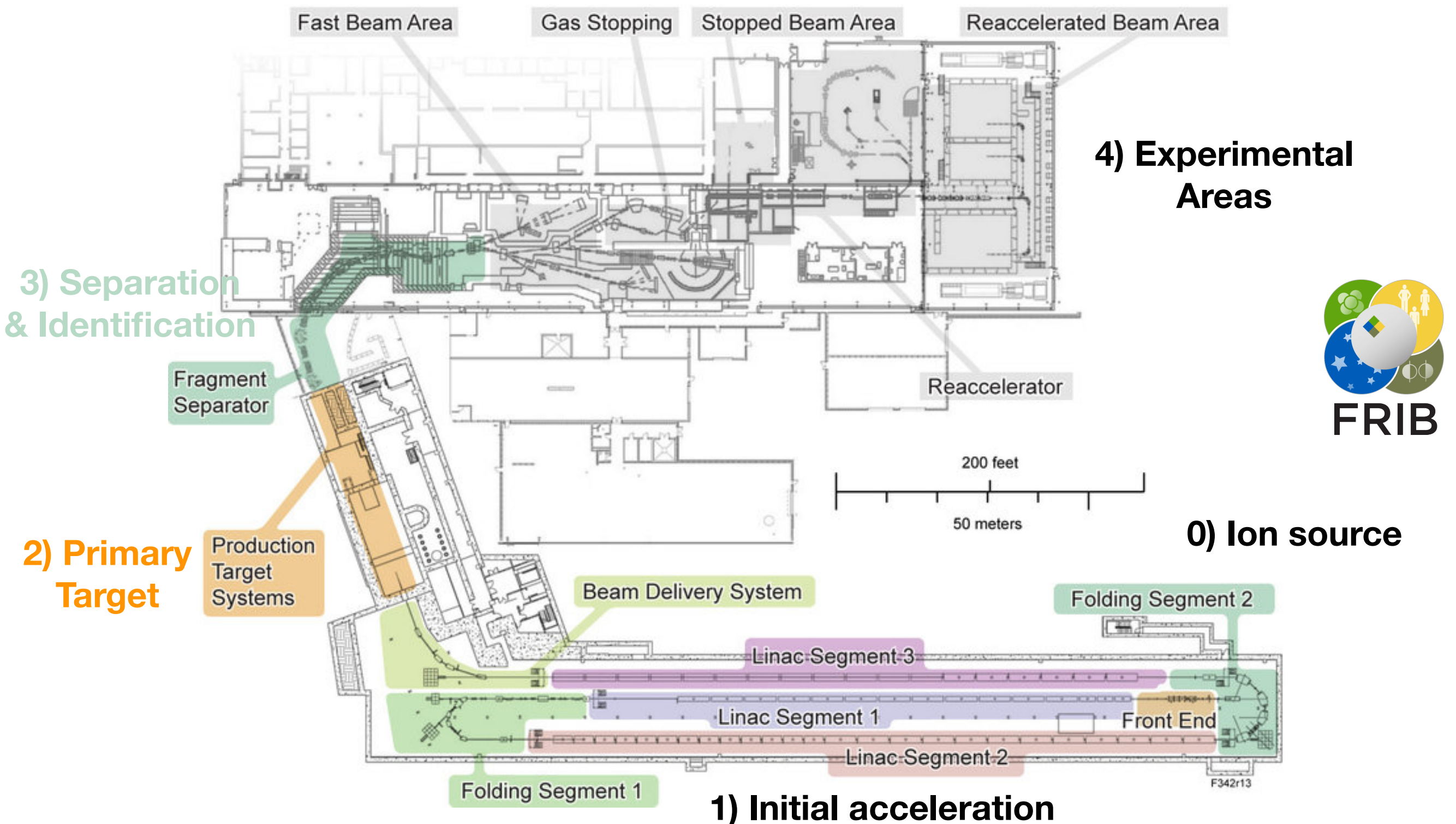


- Homogeneous electric field
- Ionization and drift of electrons towards the pads, or detection plane
- 2D picture of all charged particles in the volume from the pads + third dimension from the time of arrival

- Most nuclei we produce do NOT exist in nature
- Need an accelerator!



- FRIB = Facility for Rare Isotope Beams (started in June 2022)
- Primary beams from ^{16}O to ^{238}U from 150 MeV/u to 290 MeV/u ($\beta = 0.5\text{-}0.6$)

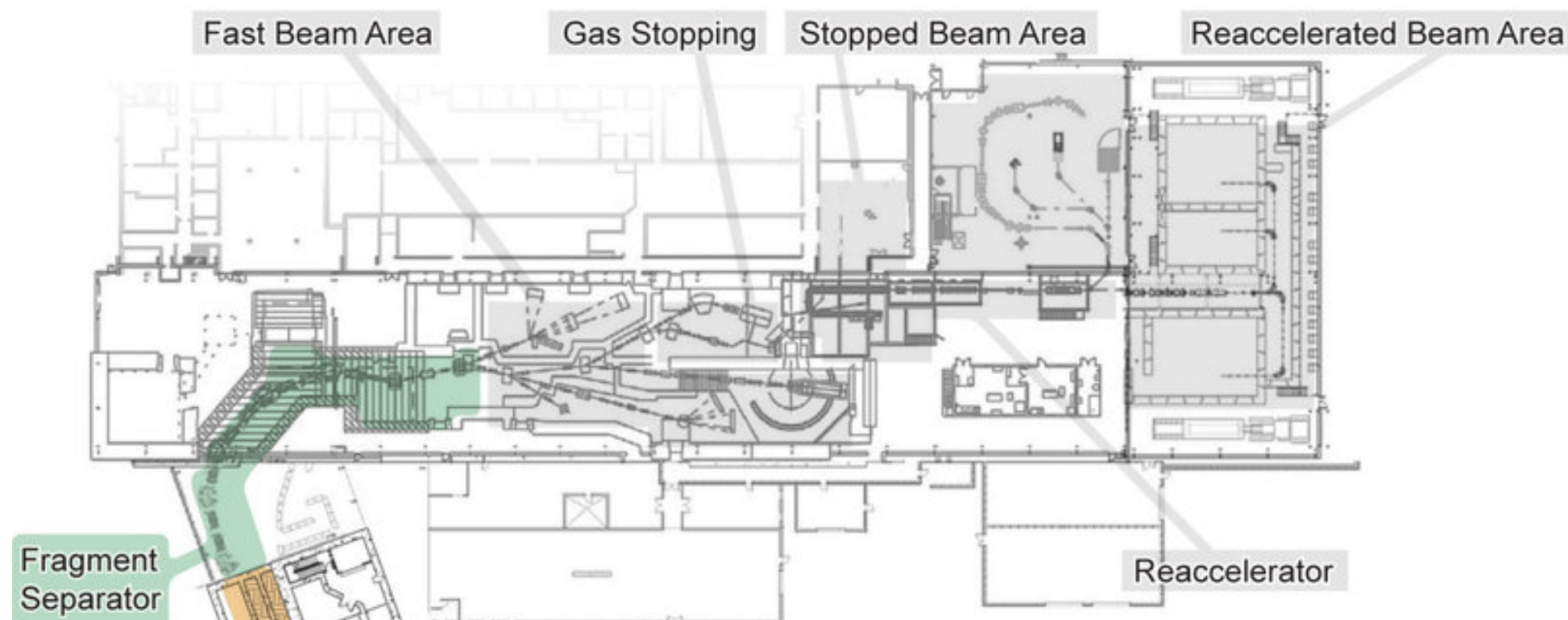




<https://www.youtube.com/watch?v=P4rG-5y9ums>

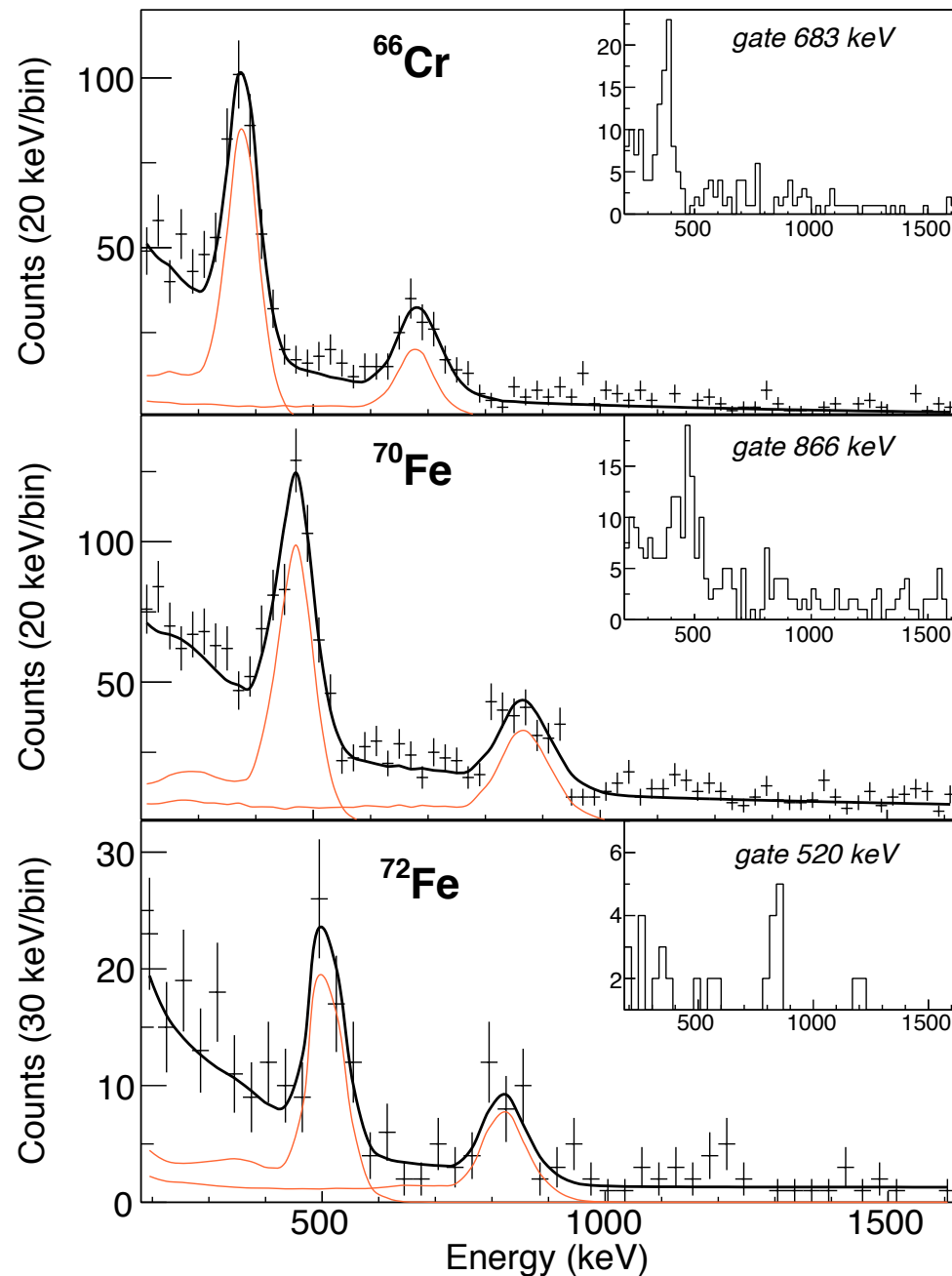
4) EXPERIMENTAL AREAS

- Use the secondary beam we created & separated before
- Collision with a secondary target
- Need to detect the products of those reactions: protons, neutrons, electrons, γ rays (photons), heavier particles...
- For every experiment, we will need to choose the detection system that fits BEST
- We don't have a system that can do everything...

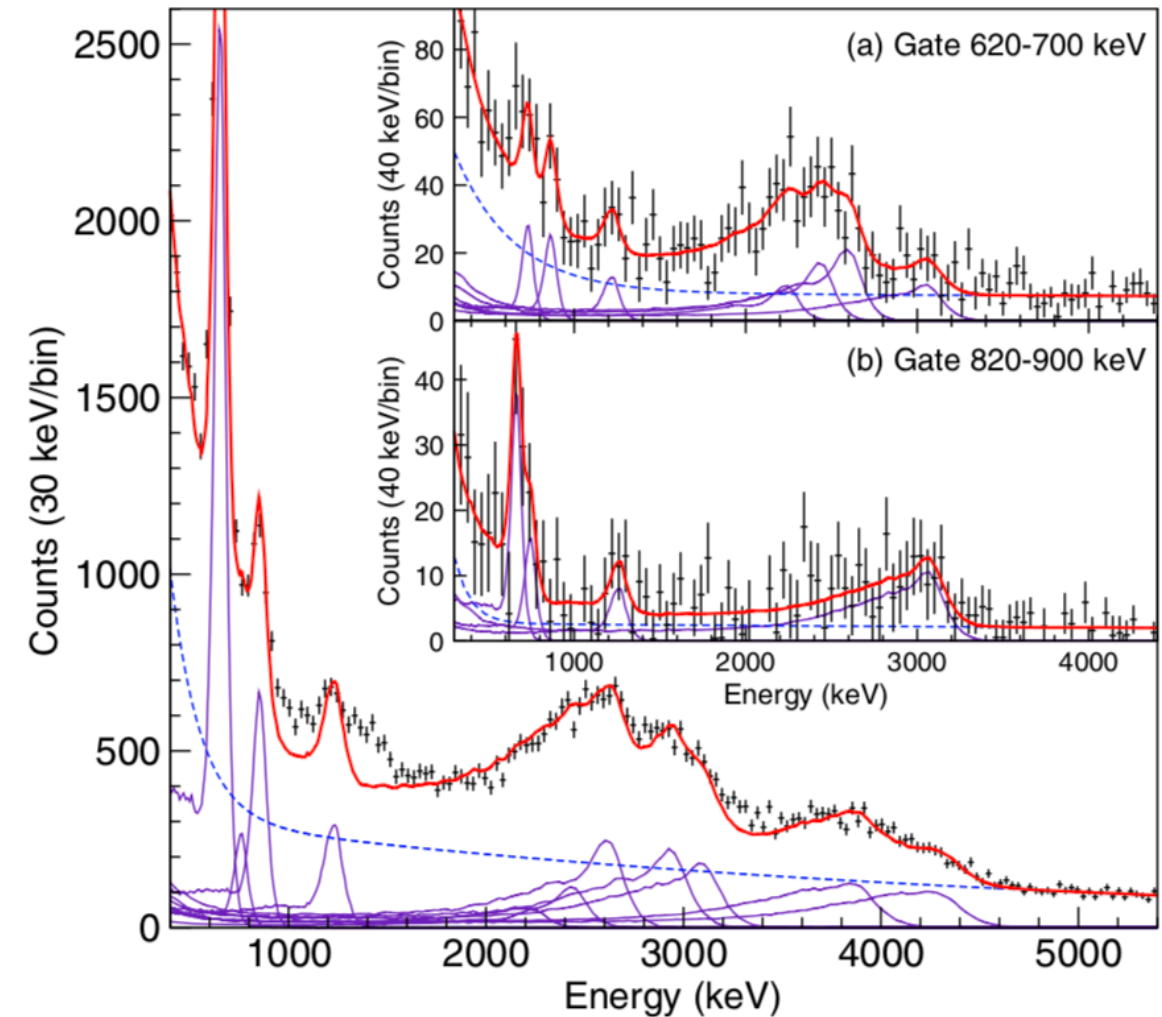


1ST SEASTAR CAMPAIGN RESULTS

C.S. *et al.*, PRL **115**, 192501 (2015)



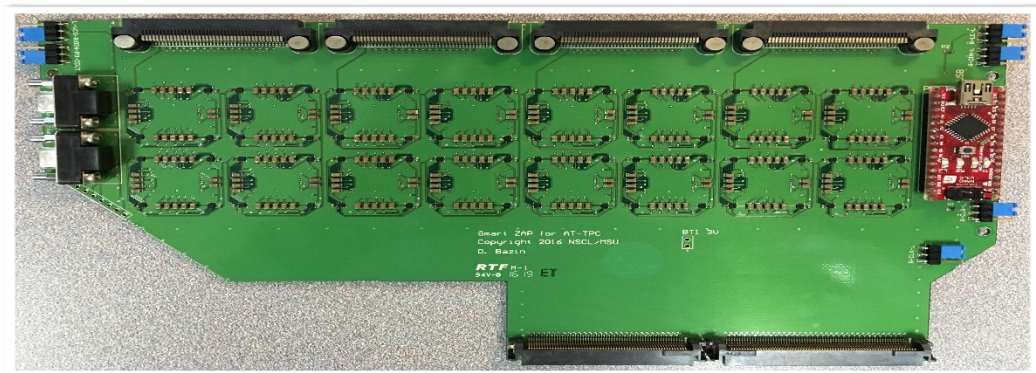
L. Olivier *et al.*, PRL **119**, 192501 (2017)



X.Y. Liu *et al.*, "Spectroscopy of $^{65,67}_{27}\text{Mn}$: Strong coupling in the $N = 40$ "island of inversion" ", Phys. Let. B **784**, 392–396 (2018).

M.L. Cortés *et al.*, "Inelastic scattering of neutron-rich Ni and Zn isotopes off a proton target", Phys. Rev. C **97**, 044315 (2018).

D. Bazin @ NSCL



AT-TPC filled with 200 Torr of He (90%) + CO₂ (10%)

